

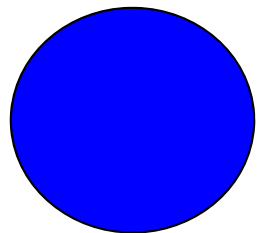
KAGRA+

GWADW
May 2018

Tokyo Institute of Technology

Kentaro Somiya
on behalf of the KAGRA collaboration

(note: KAGRA+ is not yet officially approved in the collaboration)



K.Somiya

GW detectors

USA

Europe

Japan

LIGO

Virgo

GEO

TAMA

1G

**Advanced
LIGO**

**Advanced
Virgo**

GEO-HF

KAGRA

A+

AdV+

KAGRA+

Voyager

Einstein Telescope

2022~24?

Cosmic Explorer

3G

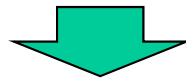
Possible KAGRA+ schedule

**2019~20:
KAGRA's first detection**

Design study of KAGRA+



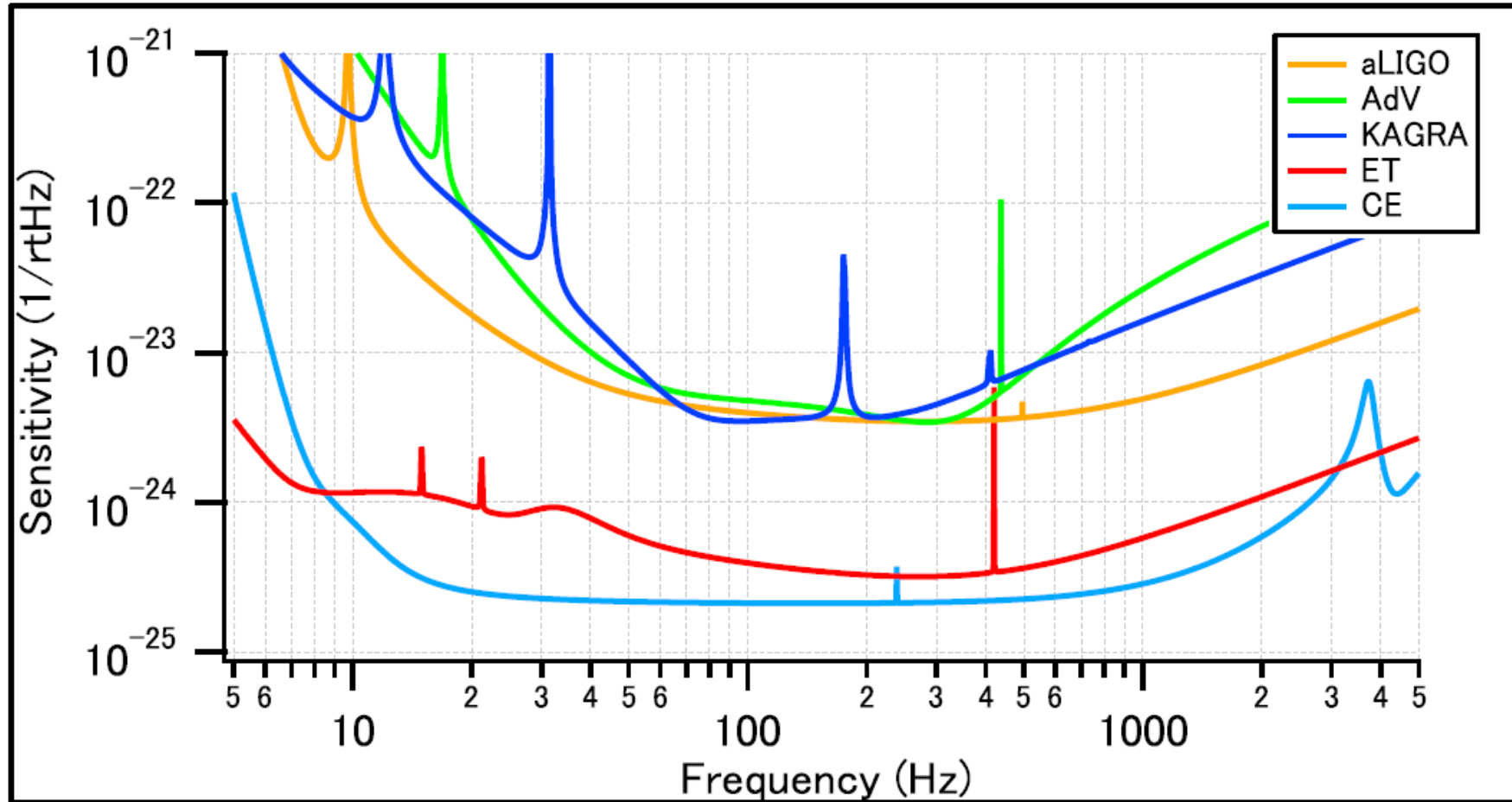
**2019 or 2020:
Application to a grant (\$5M?)**



2020~22: Implementation of new items

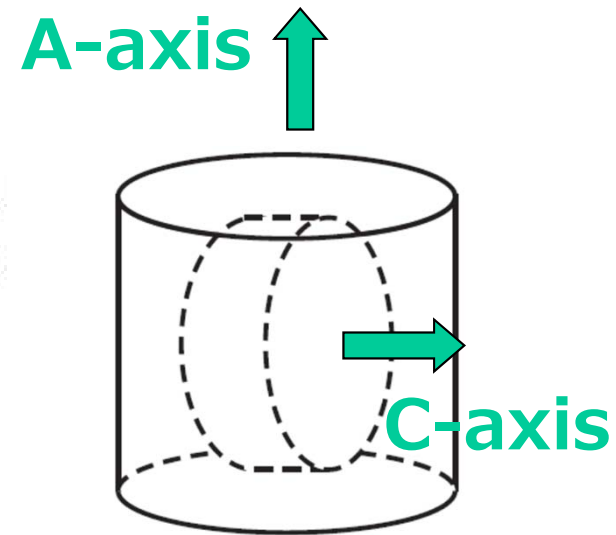
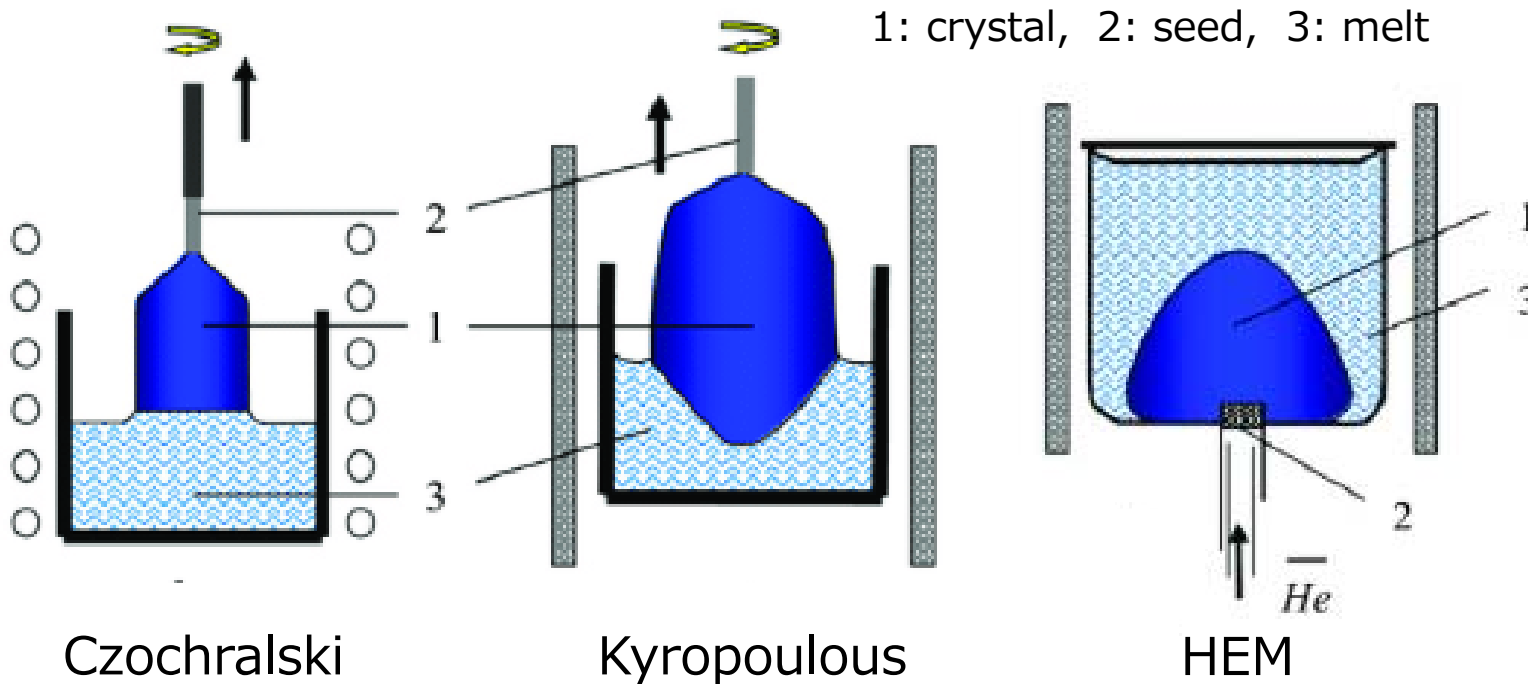
2022~24 KAGRA+

Sensitivity comparison



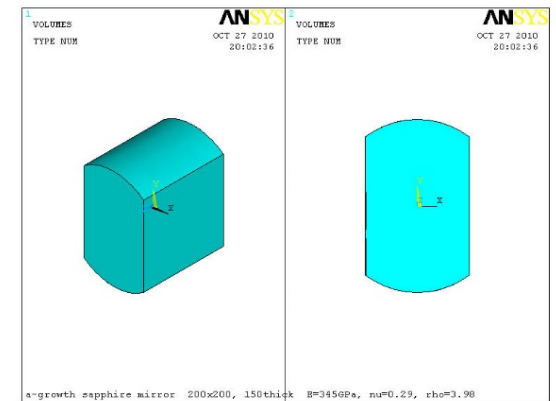
| | aLIGO | AdV | KAGRA | ET-LF | CE |
|-------------|-------|-----|-------|-------|----|
| High power | ☺ | ☺ | ☺ | | ☺ |
| Cryogenic | | | ☺ | ☺ | ☺ |
| Underground | | | ☺ | ☺ | |

Heavier Sapphire for KAGRA+?



| | crystal size | crystal axis | quality |
|-------------|--------------|----------------|---------|
| Czochralski | medium | can be c-axis | bad |
| Kyropoulous | big | usually a-axis | good |
| HEM | very big | usually a-axis | bubbles |

Reference: Kawaminami et al., *J. of the Cer. Soc. of Japan* 122, 695 (2014)



Kamaboko mirror (40kg)

Heavier Sapphire for KAGRA+?

[Kawaminami et al., J. of the Cer. Soc. of Japan 122, 695 (2014)]

- A Japanese company has developed TSMG (Top Seeded Melt Growth) method which can produce big, c-axis, good sapphire crystals.
- KAGRA's ITMs were produced with this TSMG method.
- Eiichi is to establish a new collaborative project to produce high quality larger crystals with this company, aiming at 100kg crystal(ϕ 400mm x t200mm).

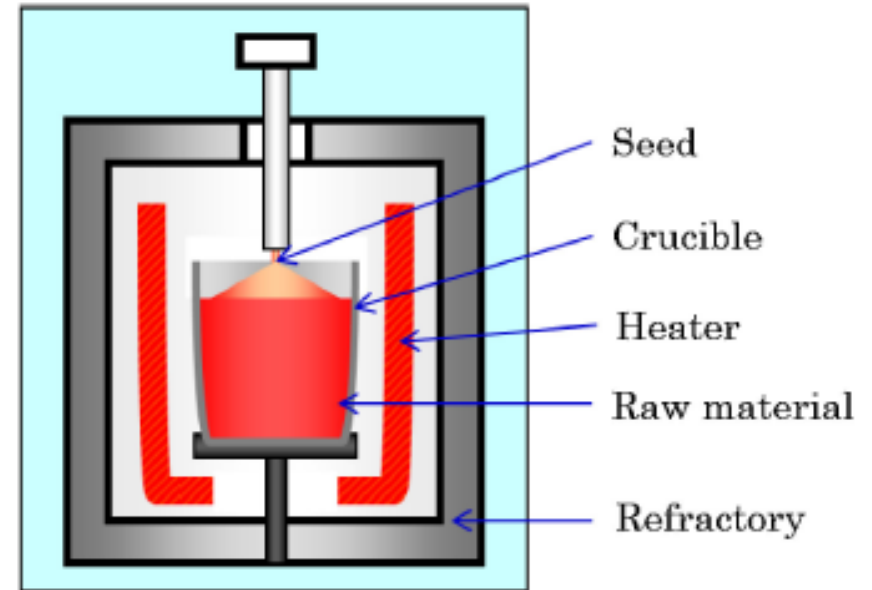
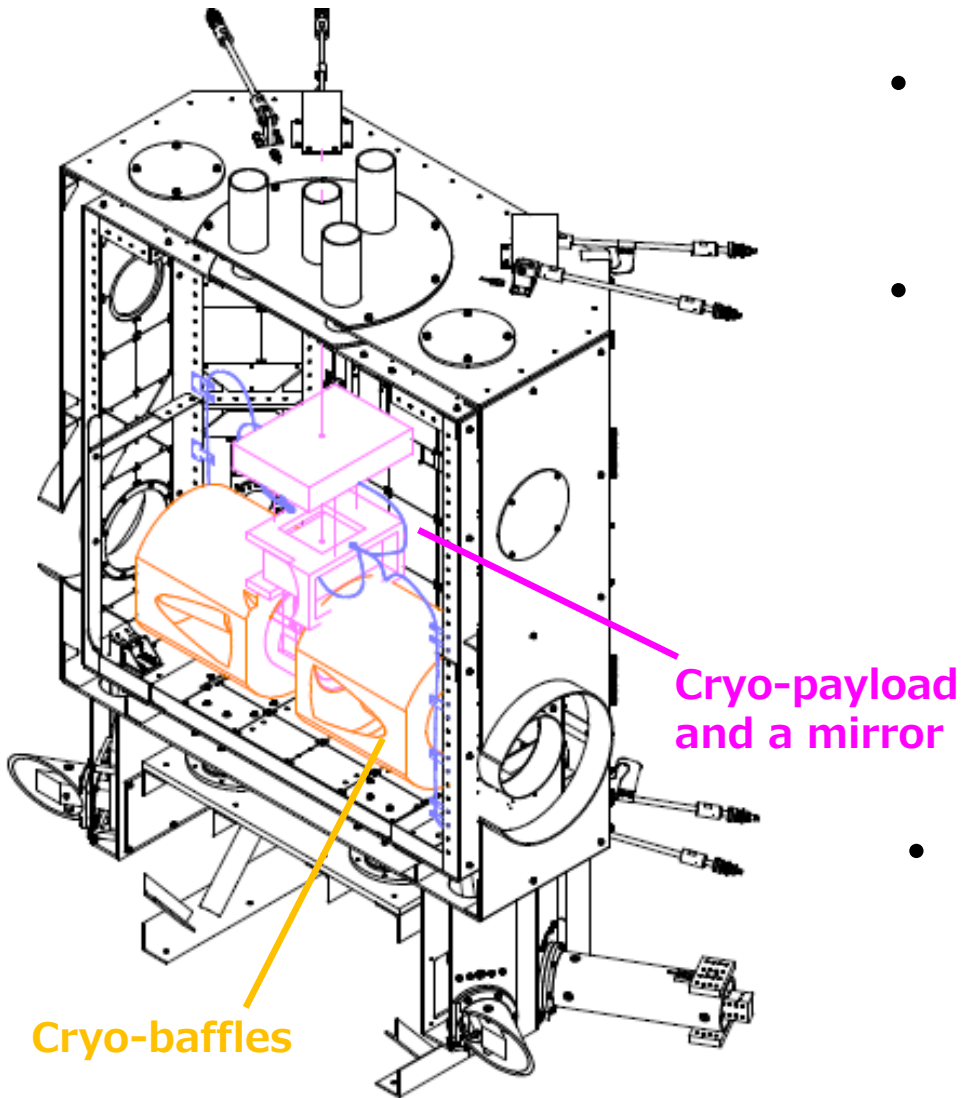


Fig. 1. Schematic illustration of the TSMG method.

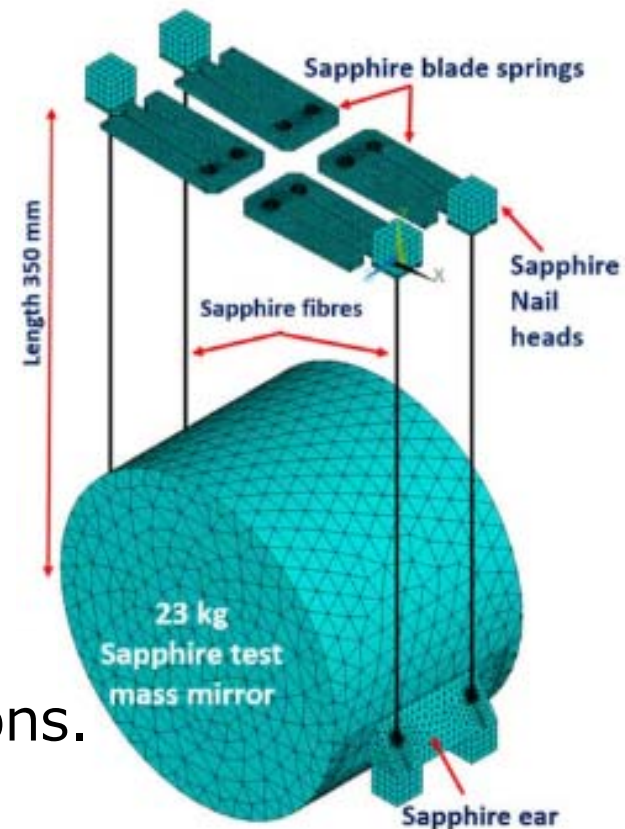
- **Having both advantages of CZ and KY**
- **Small rotation over short pulling distance**
- **Arbitrary direction of growth**
- **Oxygen vacancies (UV absorption)**

Additional issues to be solved

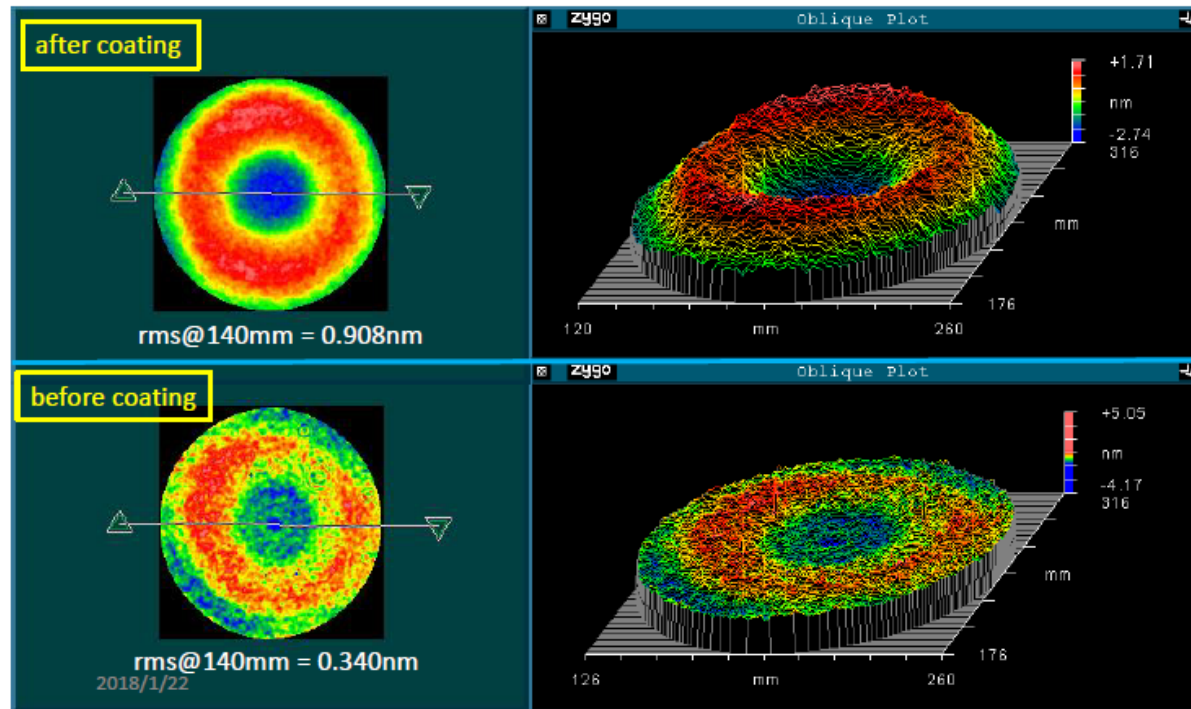


- A cryostat is quite full with the current $\phi 220 \times t150$ crystal and cryo-baffles.
- Kazuhiro Yamamoto says x1.2 may be ok but x1.5 larger mass would be hard. ($23\text{kg} \times 1.2^3 = 40\text{kg}$)

- Load capacity of the blade and bondings are also an issue.
- Type-A suspension may need modifications.



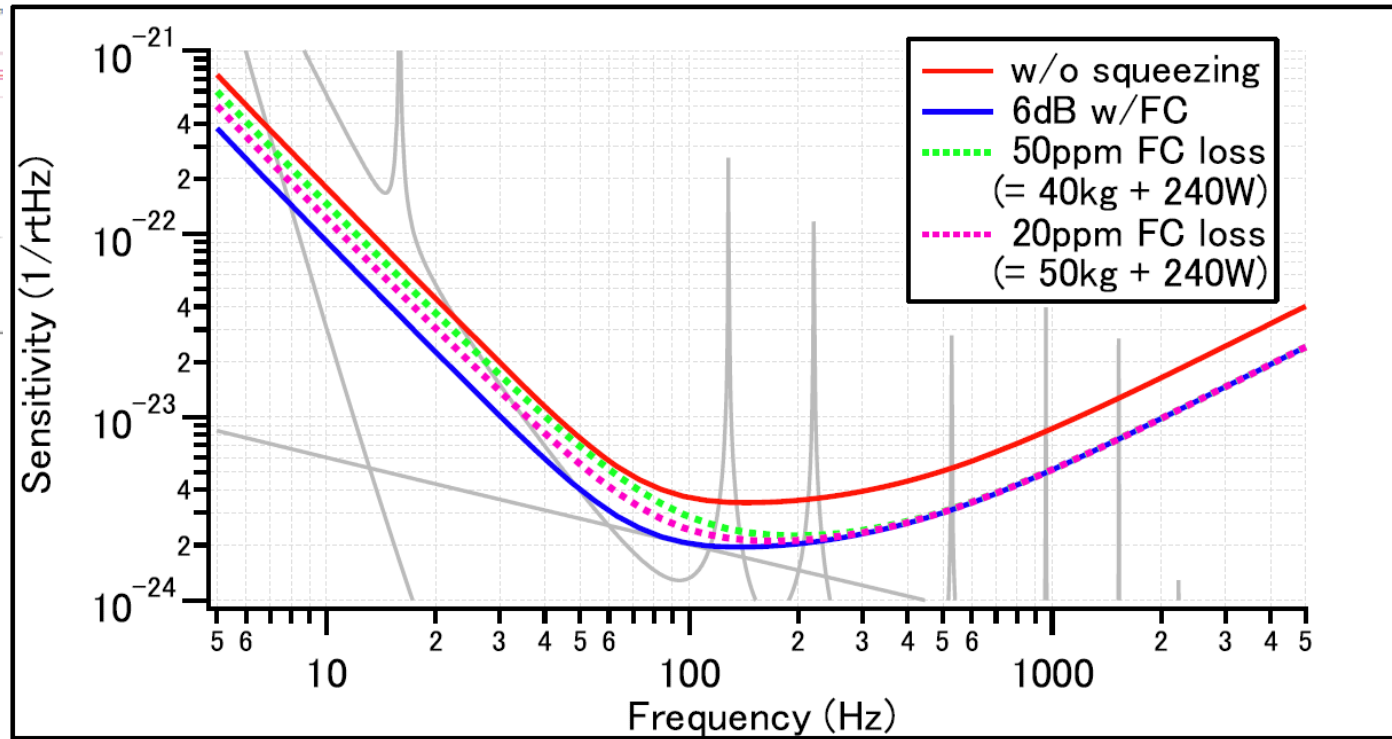
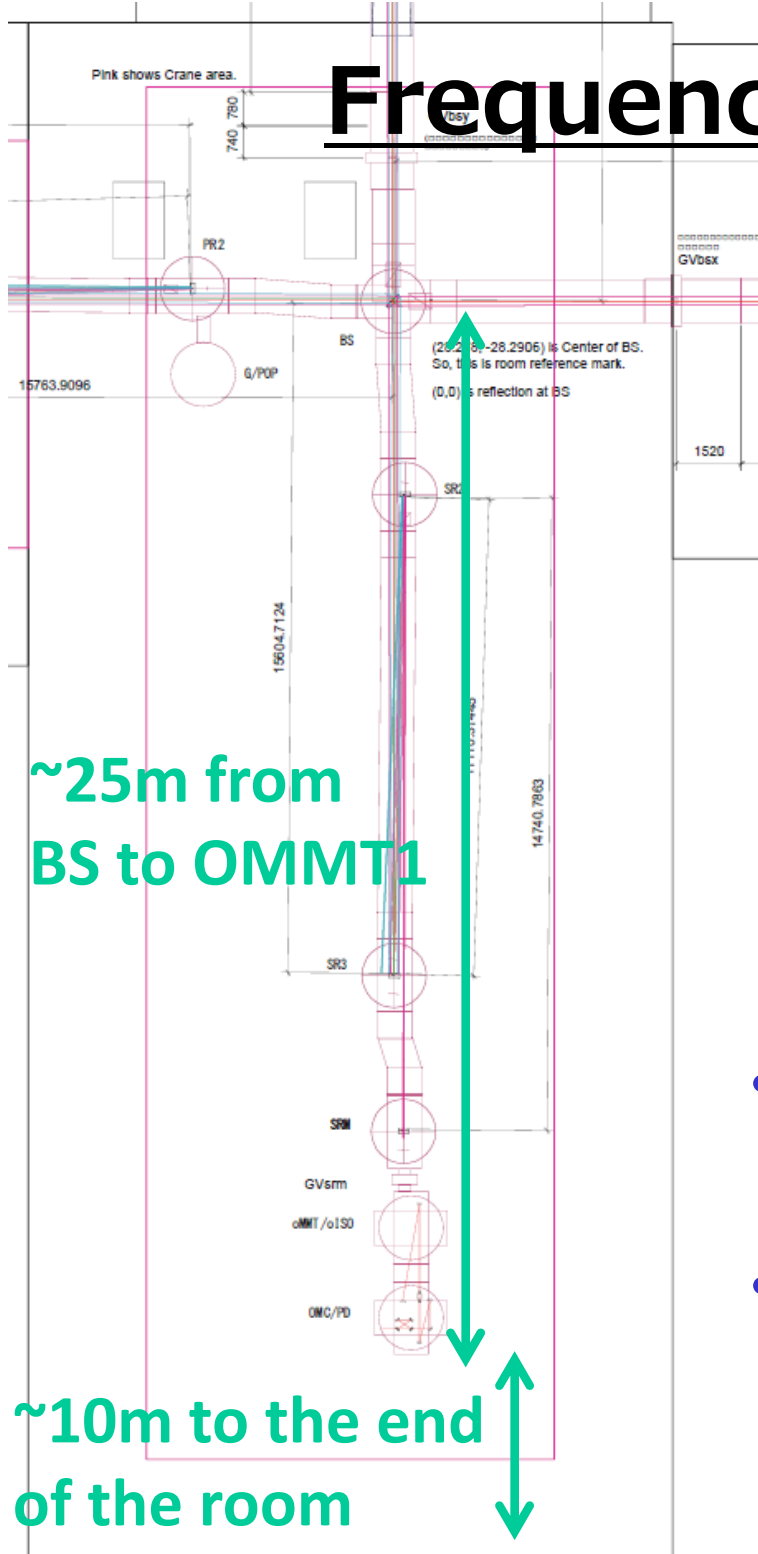
Cost of sapphire mirrors



- crystal: ~\$100k per piece
- polish: ~\$400k per piece
- coating: ~\$100k per piece

Most of the budget for the upgrade would be spent for the mirrors.

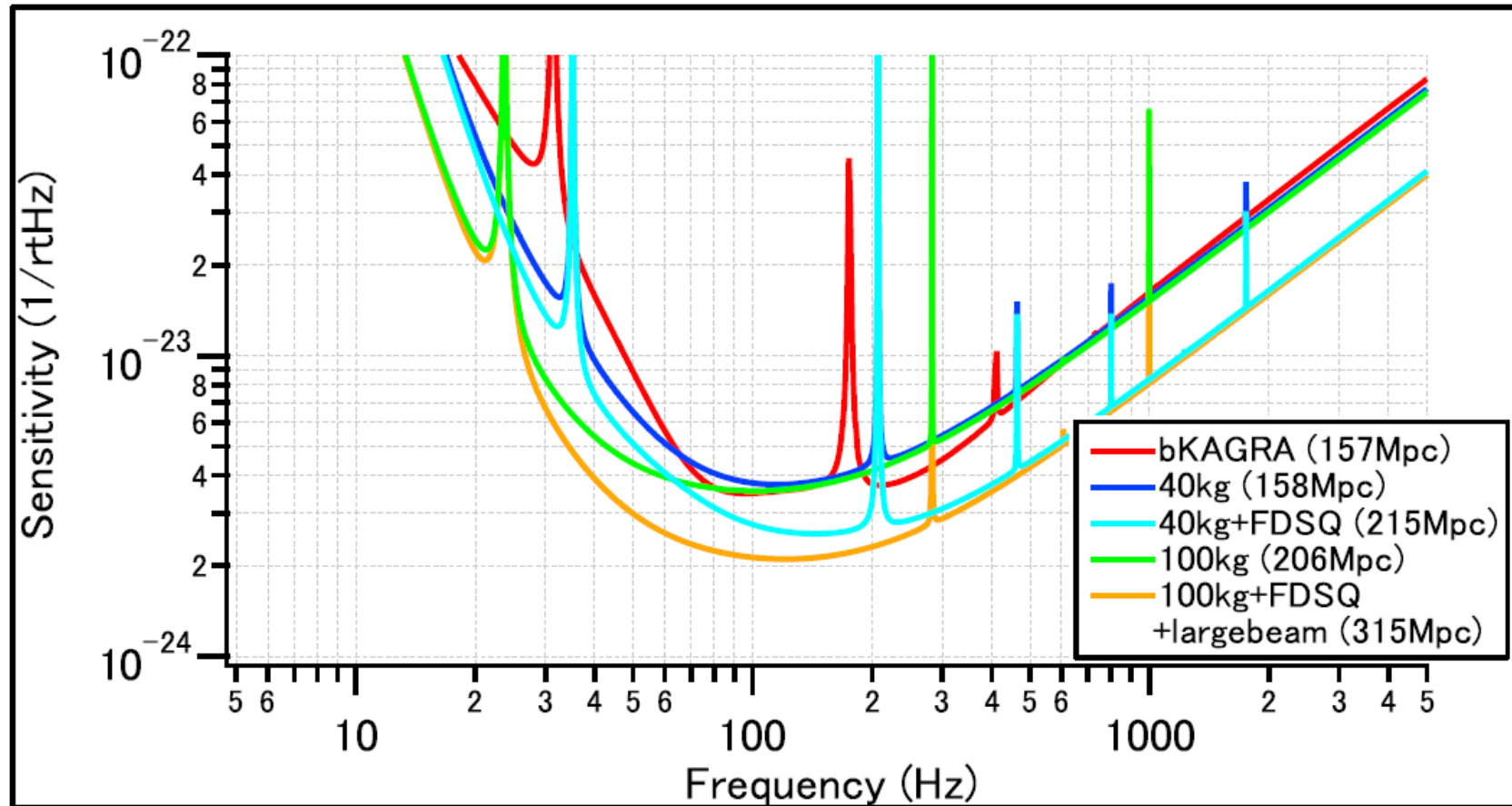
Frequency-dependent squeezing



- A use of FD SQ is equivalent to a use of larger test masses (for QN).
- R&D with a 300m FC has been performed in TAMA.

[see Eleonora's talk on Wed]

Sensitivity curves with heavier masses



Required items

158Mpc: Mirror

\$3M

215Mpc: Mirror + FDSQ

\$6M

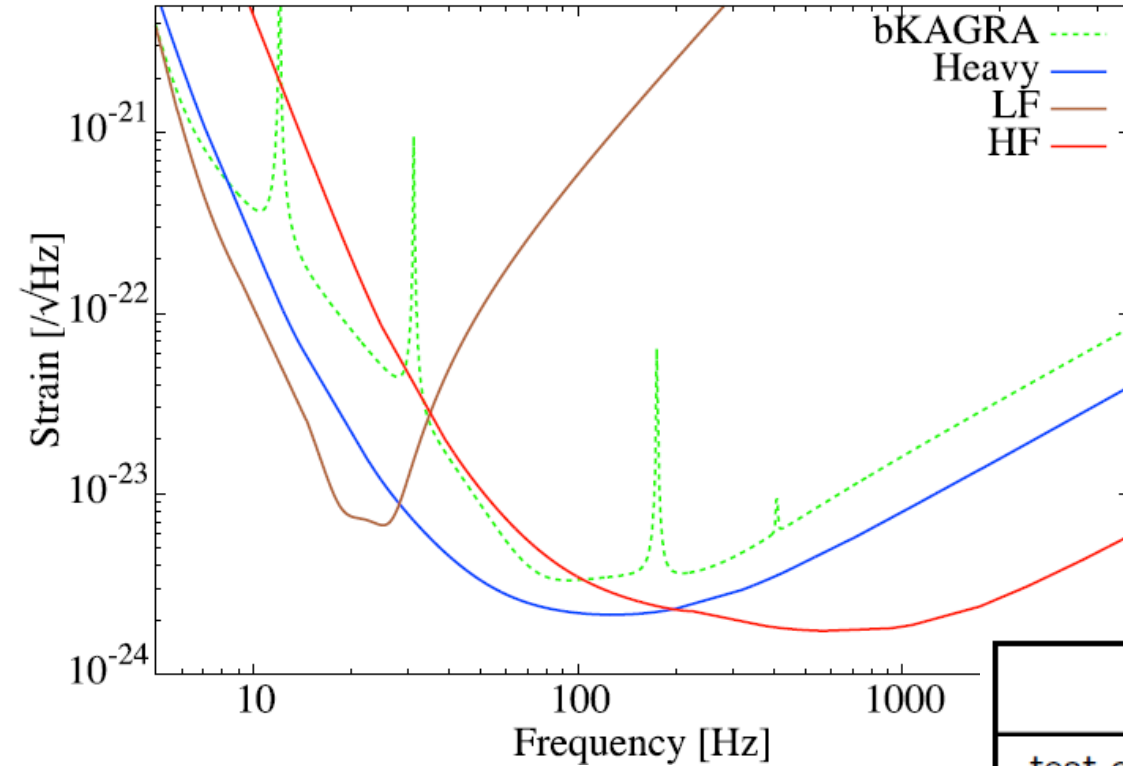
206Mpc: Mirror + Cryostat/payload + Type-A

\$8M

315Mpc: Mirror + Cryostat/payload + Type-A + FDSQ + RMs

\$12M

Other options



LF: low power + longer suspension
 HF: high power + thicker suspension

Theory people in Japan compared the science cases to find that HF is as promising as Heavy.

[Nishizawa]

| | bKAGRA | LF | Heavy | HF |
|--|--------|----|-------|----|
| test of GR with BH ringdown | × | × | △ | ○ |
| existence of IMBH from hierarchical growth | △ | △ | ○ | △ |
| existence of stellar-mass BBH from popIII | × | × | × | × |
| sky localization for BBH (identifying host galaxy) | △ | × | ○ | ○ |
| pulsar ellipticity | × | × | △ | ○ |
| NS equation of state | × | × | △ | ○ |

Heat extraction with thick fibers

[Khalaidovski 2014]

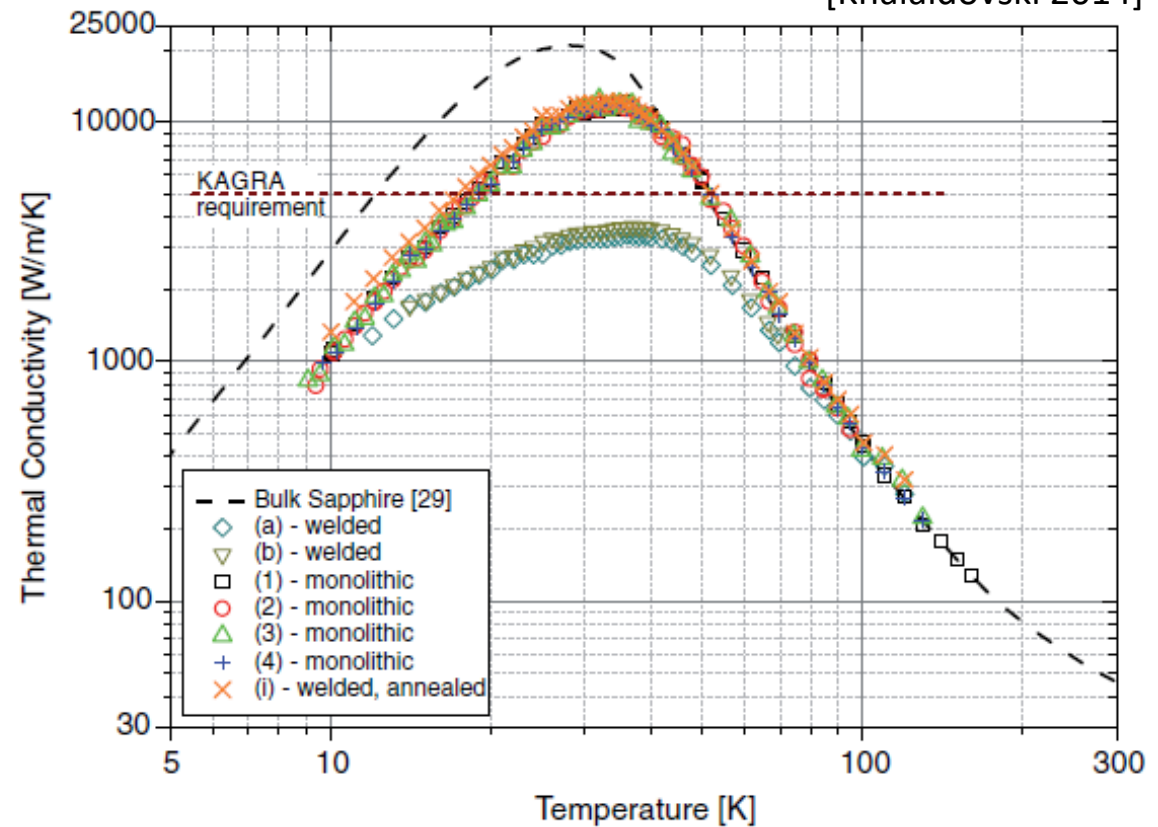
Removable heat is given as

$$K_{abs} = \int_{T_1}^{T_2} \frac{\pi d^2 \kappa(T)}{l_{sus}} dT$$

where T_1 (IM) is 16K.

Sapphire thermal conductivity is approximated to be

$$\kappa = 5270 d \times T^{2.24} \text{ (W/m/K)}$$



| | | | | | | | |
|--------------|----------------|----------------|--------------|----------------|----------------|----------------|----------------|
| T=22K | d=1.6mm | d=2.4mm | T=24K | d=1.6mm | d=2.4mm | T=25.9K | d=2.6mm |
| l=35cm | 0.86W | 2.9W | l=35cm | 1.3W | 4.4W | l=35cm | 7.7W |
| l=20cm | 1.5W | 5.1W | l=20cm | 2.3W | 7.7W | T=38.5K | d=1.6mm |
| | | | | | | l=35cm | 7.7W |

x10 power is capable.

High-power laser and PRG



Currently we have two 40W fiber lasers. Our solid-state amp turns out to be not easy to handle.



| arm loss | contrast | PRM | PRG |
|----------|----------|-----|-----|
| 45ppm | 99.9% | 90% | 10 |
| 30ppm | 99.9% | 93% | 15 |
| 25ppm | 99.9% | 94% | 17 |

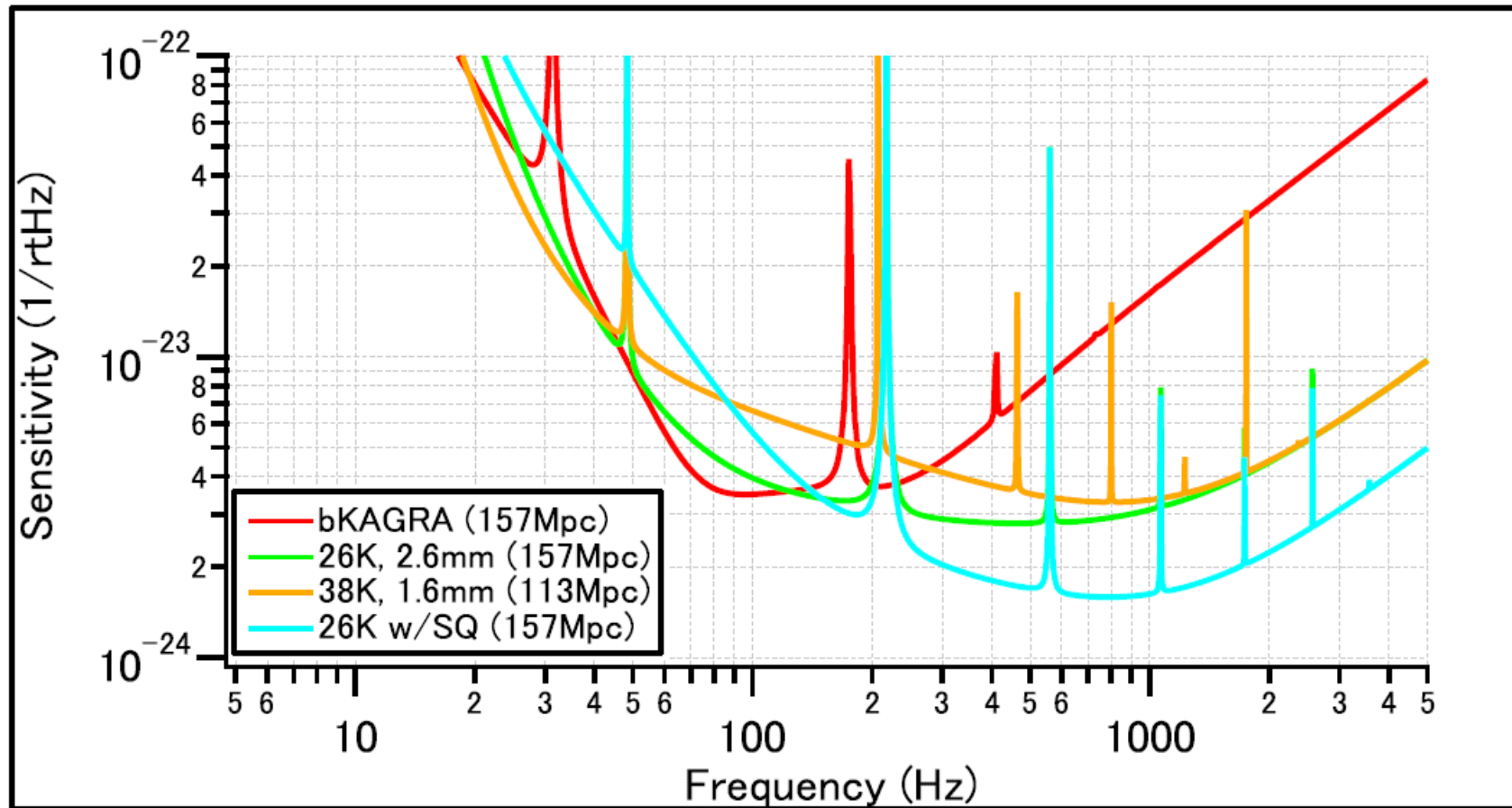
PRG could be as high as 17 if the arm loss turns out to be small.



Other issues of high power

1. Thermal lensing in BS
2. Input Optics chain (IFI, EOM, IMC, etc.)
3. Output Optics (BD, shutter, PD, etc.)
4. PI

Sensitivity curves with high power



Required items

38K: 400W laser + IO (+ TCS)

\$3M

26K: 400W laser + IO + Fiber (+ TCS)

\$4M

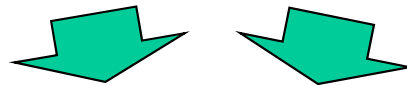
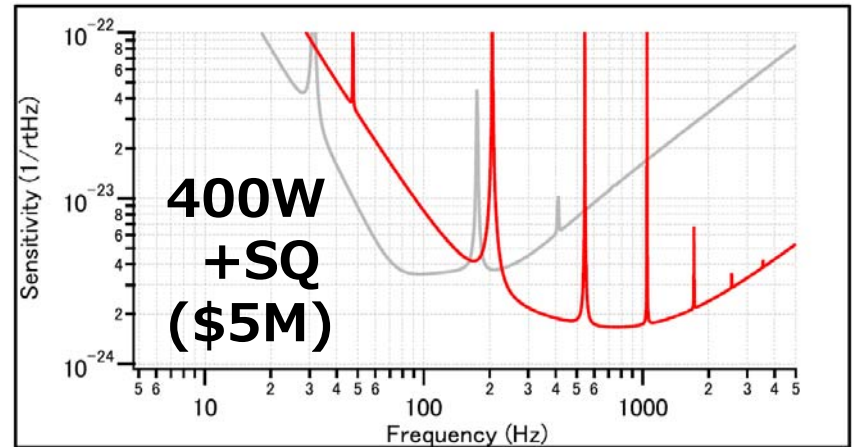
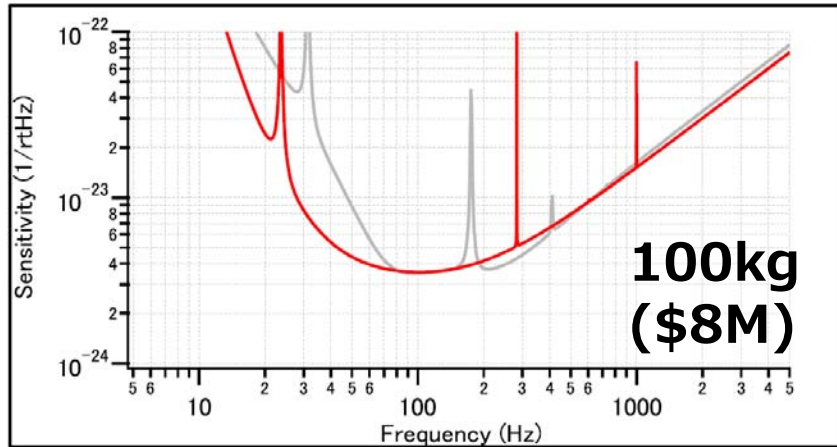
26K+SQ: 400W laser + IO + Fiber + SQ (+ TCS)

\$5M

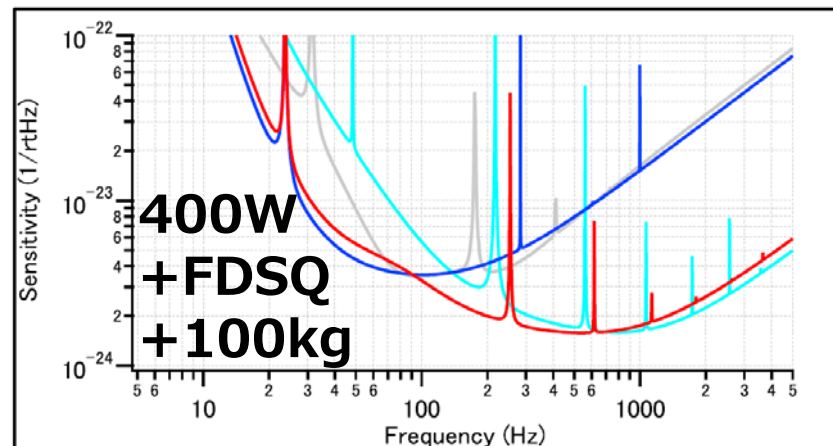
KAGRA+ possible chart

bKAGRA commissioning done

either  or  in 2022~24 (158Mpc)



combine them in ~2028? (293Mpc)



KAGRA++?

Summary

- 1. Some people in KAGRA have started working on near-future upgrade plans (“KAGRA+”).**
- 2. Possible upgrades would be (i) high-power or (ii) larger mass. An alternative may be low-f.**
- 3. If we are aiming at an observation in 2022~24, the schedule is quite tight.**
- 4. I would suggest that we choose one of (i)(ii) for KAGRA+ and keep the other for the further future.**
- 5. For a detailed optimization, Yuta will explain a cool method in the next talk.**