



# *Overview on thermal noise reduction research for future GW detectors*

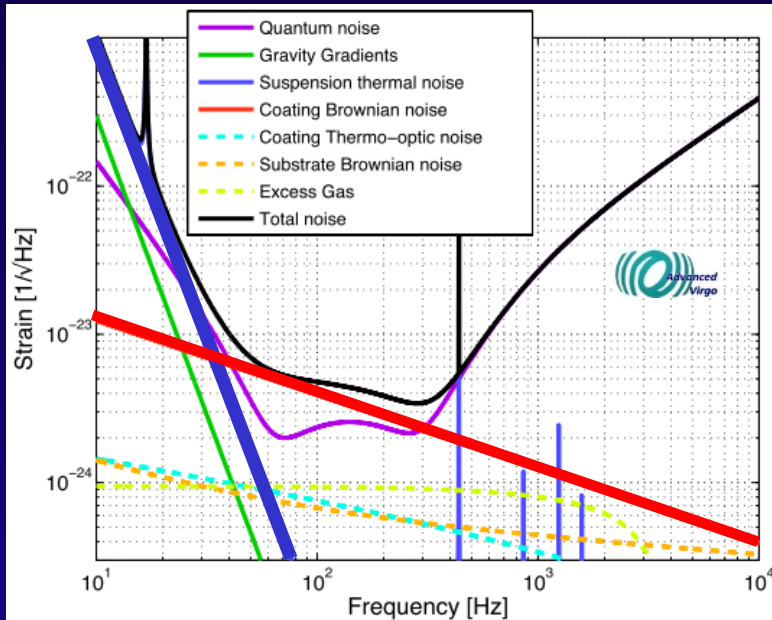


G. Cagnoli, LMA

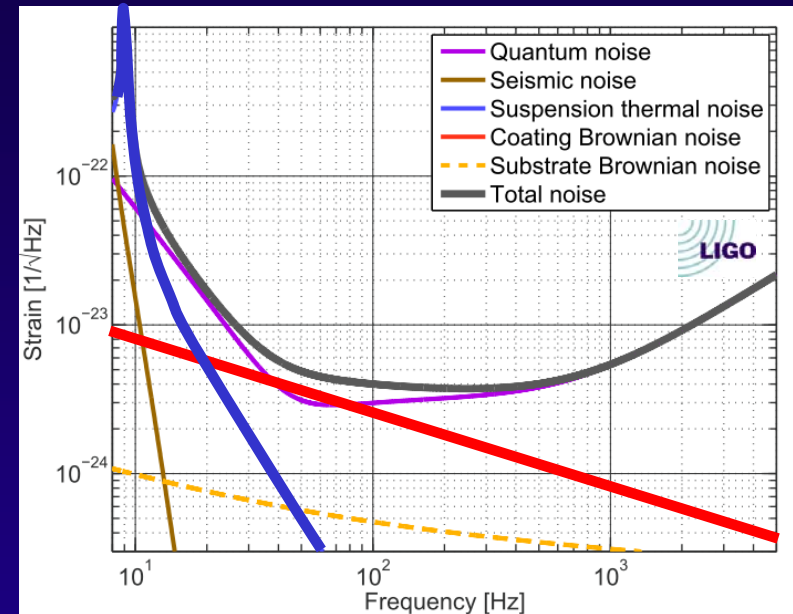


GWADW 2014  
Takayama – 26 May 2014

# Sensitivity limits in Advanced Detectors

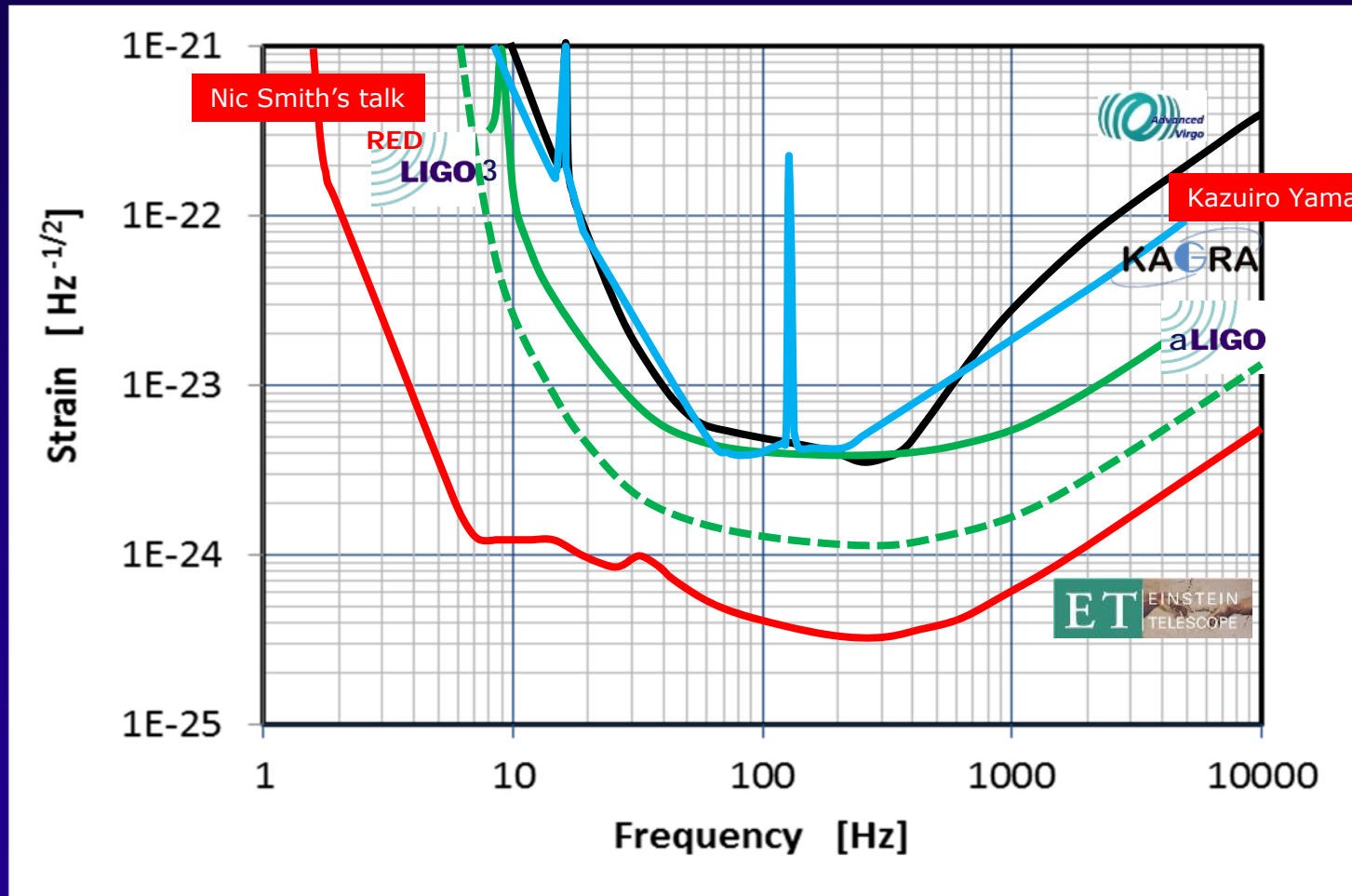


- **Suspensions:**
  - Silica fibres
  - Length <1m
  - Thermoelastic cancellation
  - Vertical mode frequency <10 Hz

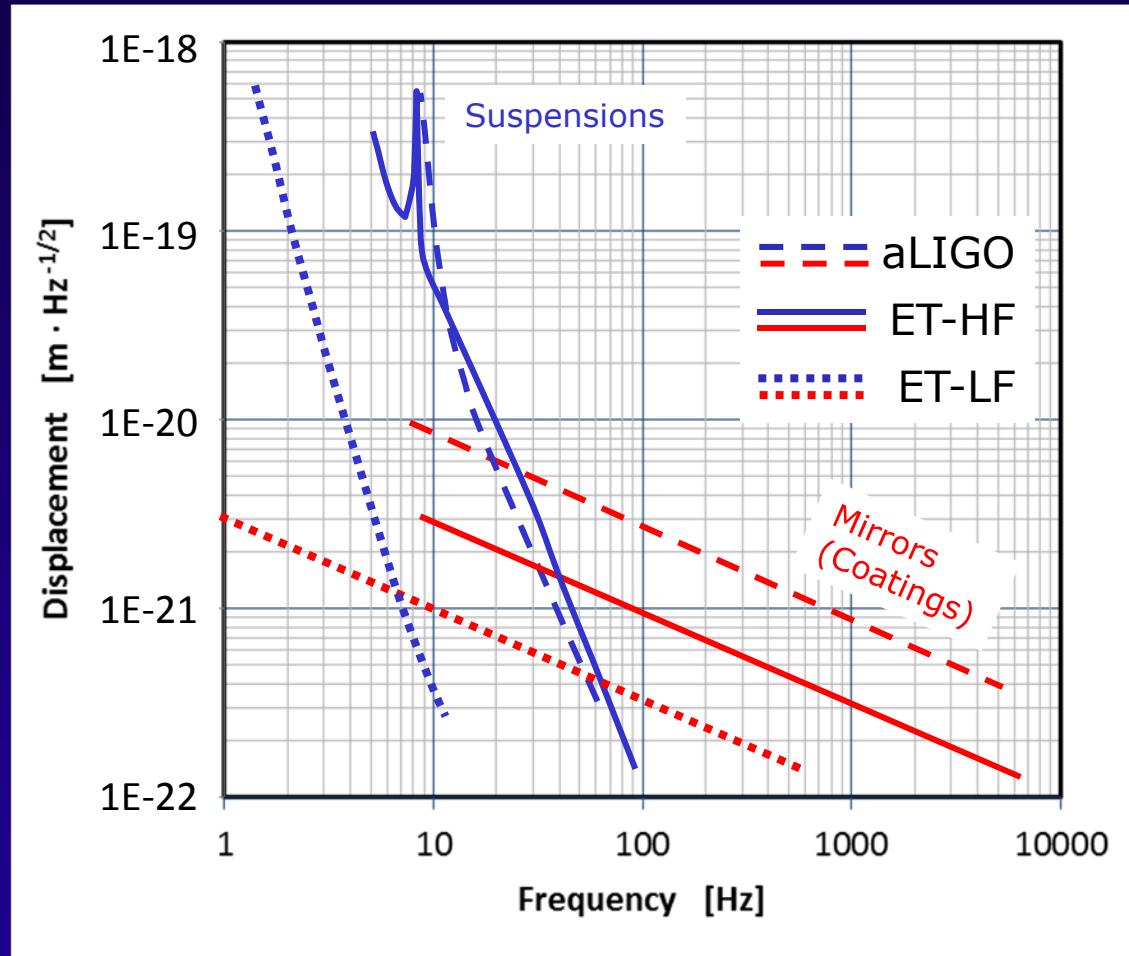


- **Substrates and Coatings:**
  - Silica substrates
  - Optimal aspect ratio
  - Ti:Ta<sub>2</sub>O<sub>5</sub>/SiO<sub>2</sub> coatings
  - Optimal coating design

# Sensitivity curves beyond the Advanced detectors



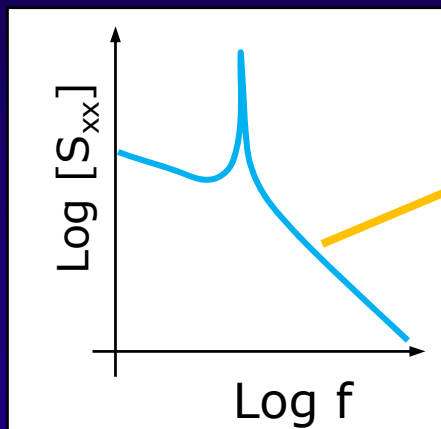
# Comparison of thermal noise levels



- How do we go from Advanced to Future detectors?

# Suspension Thermal Noise

- In the detection bandwidth, interferometers work above the pendulum resonant frequency

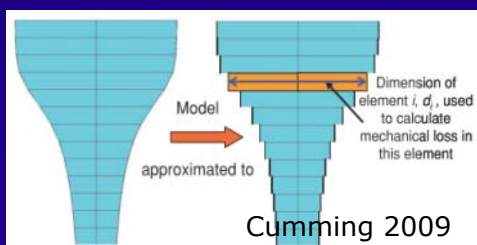


$$S_{xx} \propto T \cdot \frac{1}{m f^5} \cdot \frac{g}{L} \cdot \text{Dil} \cdot \varphi_{\text{fibre}}$$

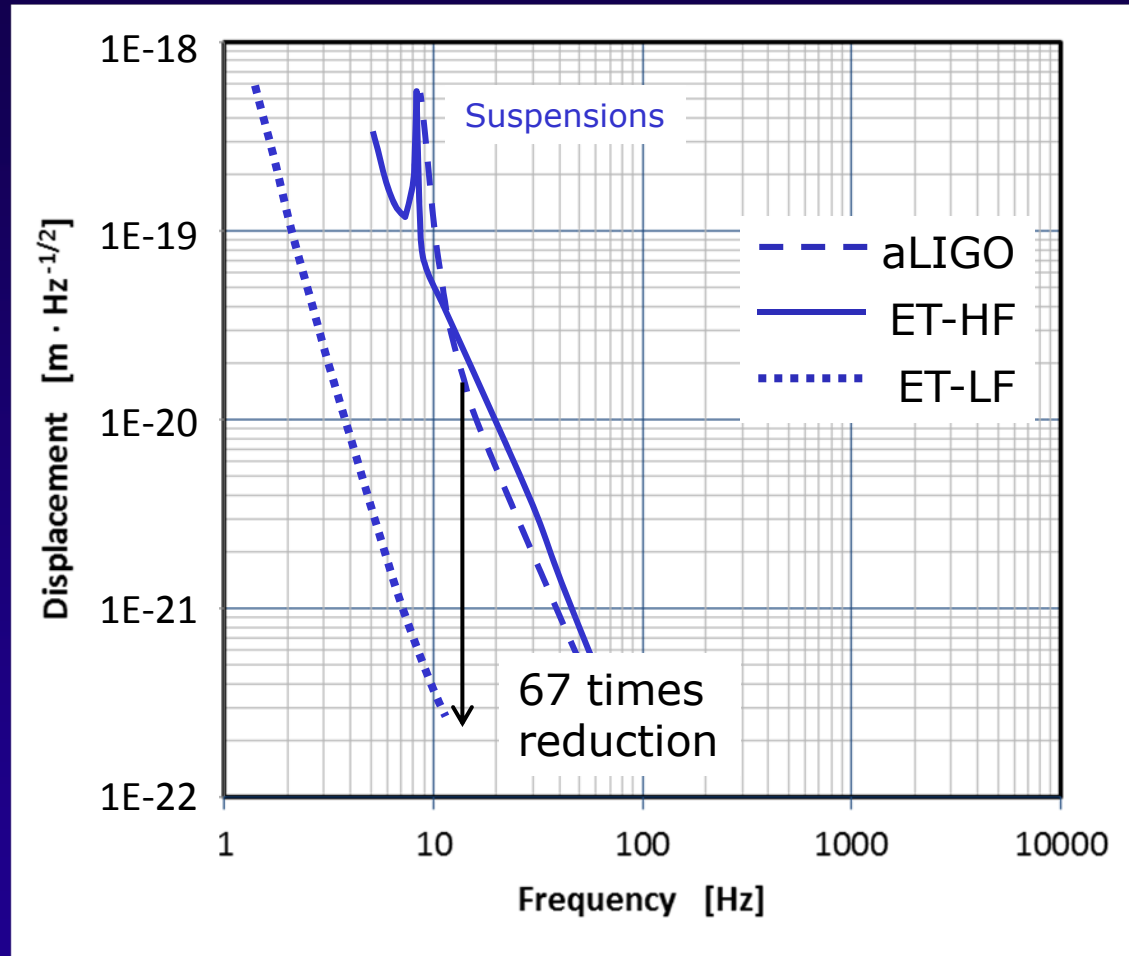
$$\text{Dil} = \frac{r^2}{L} \sqrt{\frac{\pi Y}{m g}}$$

- For fibres with a neck the dilution is almost independent on mass
- Fixing the fibres radius
  - Thermo-elastic cancellation in silica ( $\sim 190\text{Mpa}$ )
  - Sufficient heat conduction for sapphire and Si
  - Dil in ET-LF =  $2.6 \times$  Dil in aLIGO
- Summing them up:
  - $1/30$  from  $T$  –  $1/3$  from  $L$
  - $1/5$  from  $m$  –  $1/10$  from  $\varphi$

$\sqrt{S_{xx}}$  may decrease 67 times



# Comparison of thermal noise levels

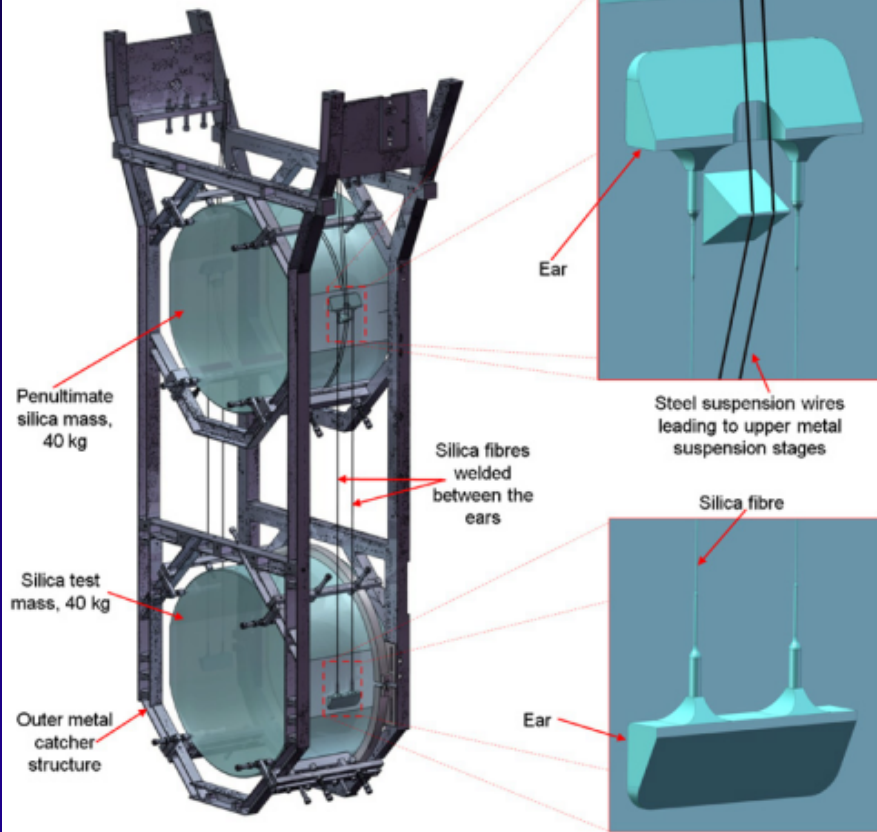


- What scientific and technological advancements are needed to secure the result?

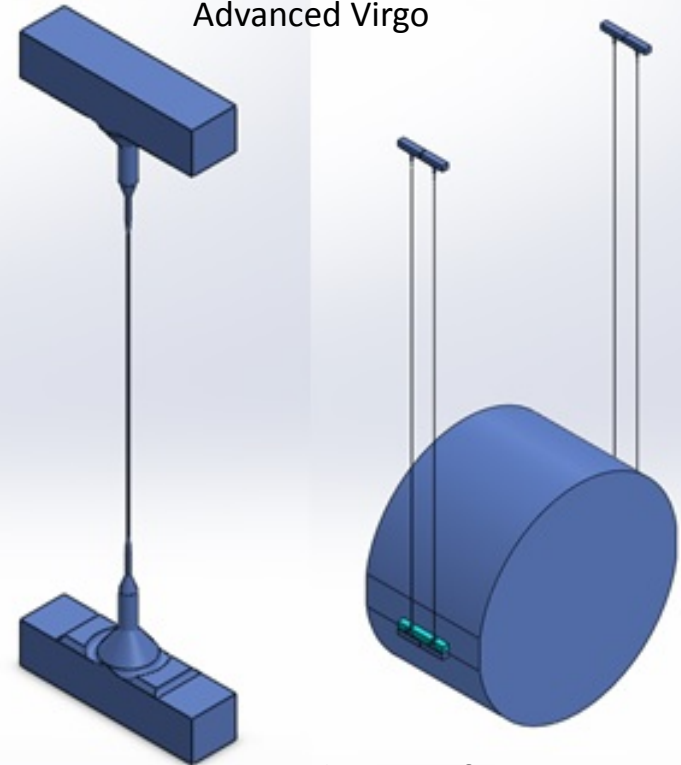
# Two concepts for warm suspensions

Cumming 2012

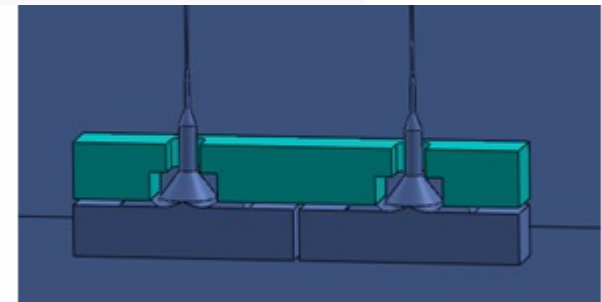
Advanced LIGO



Advanced Virgo



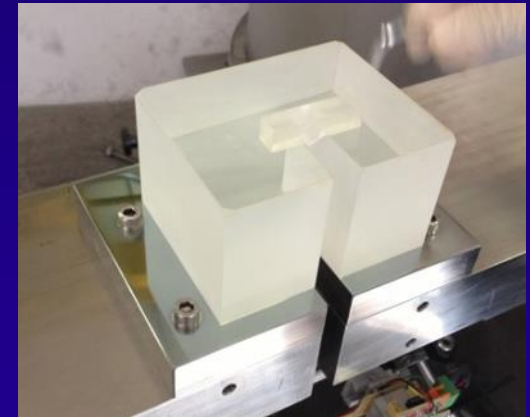
Courtesy of H. Vocca, Perugia



# Technologies for warm suspensions

- Fused silica fibres
  - Heavier masses require thicker fibres
  - Breaking strength is related to the probability to have a flaw in the fibre, which is proportional to the volume
- Tapering
  - A sharp neck increases the dilution factor
- Bonding
  - Shear or compressive stress
  - Bonding silica to silica coated steel
- Cantilever blades
  - DLC protective coating increases handling

Ross Birney's talk



Courtesy of H. Vocca, Perugia



# Technologies for cold suspensions - 1

- Crystalline fibres

- $\mu$ pulling
- Vertical growth
- Machining



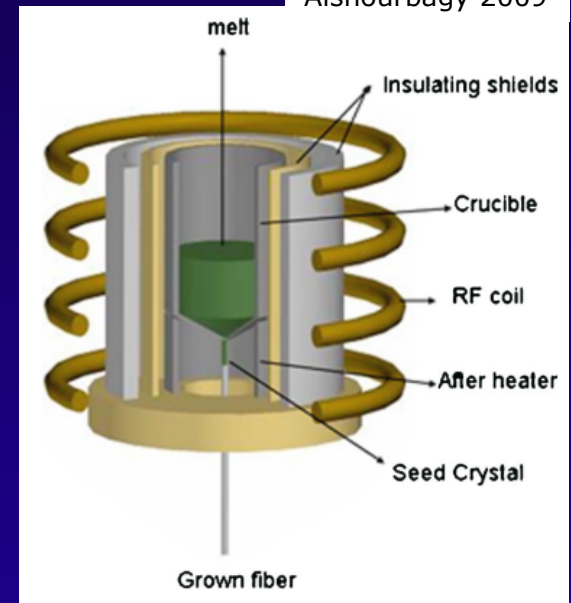
Alshourbagy 2009

- Bonding

M. Van Veggel's talk and poster

- New challenges come from:
  - 1) the cryogenic environment
    - Thermal cycling

- 2) Sapphire and silicon bonding



Sapphire bonded samples which broke across the bond surface, but with damage to the bulk sapphire. **Strengths 60-80MPa at 77K**

S. Rowan's talk, GWADW 2013

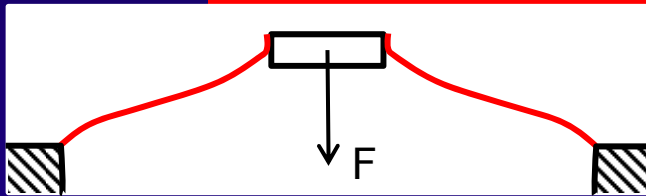
- 3) Indium bonding?

# Technologies for cold suspensions - 2

## • Crystalline LF vertical suspensions

- Cantilevers
- Elements under compression

Eric Hennes's posters



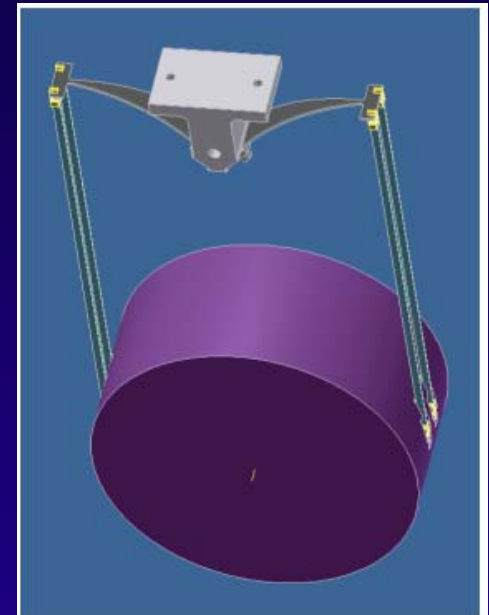
G. Hammond's poster on breaking stress of sapphire

## • Cradles

Raul Kumar's talk



A. Cumming's talk,  
5<sup>th</sup> ET Symposium



E. Majorana's talk,  
GWADW 2013

Mode	Frequency no spring Hz	Frequency spring Hz
Pendulum	0.89	0.89
Vertical bounce	106	15.3
Violin	221	221

# Scientific investigations

- Thermal noise under heat currents

Giles Hammond's poster on Johnson noise

Daniel Heinert's poster on non-eq. Th.Ns.

- Surface relaxations in mechanical losses of amorphous and crystalline materials

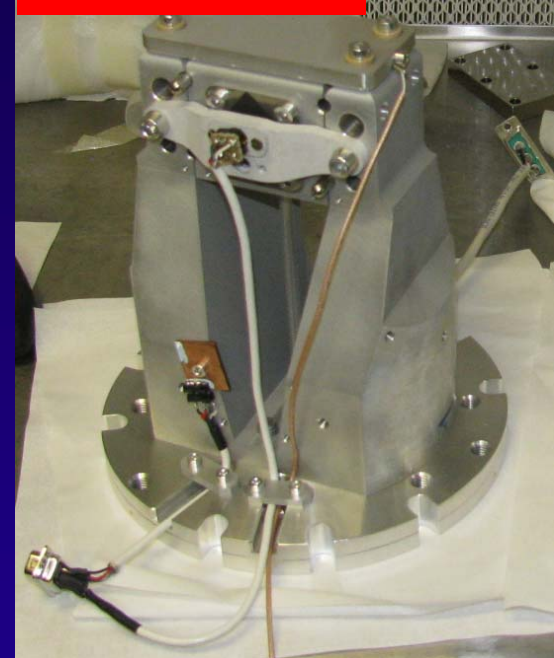
Dmitry Koptov's poster on the ESD effect on test masses

Zach Korth's poster on silicon ribbons noise

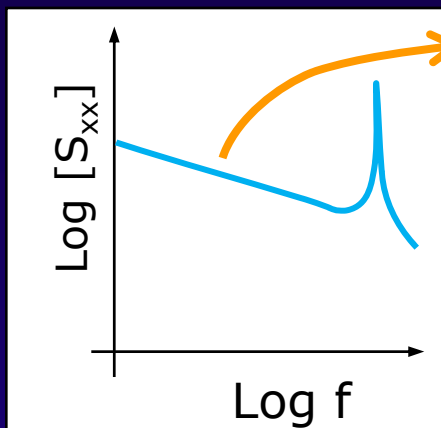
Nic Smith's poster on continuous measurement of  $Q$

- Can we have a factor 10 gain over fused silica losses?

Livia Conti's talk



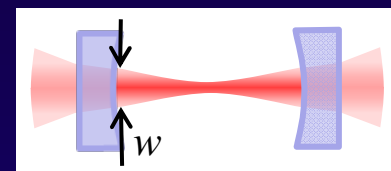
# Mirror thermal noise



Mirror noises are always below resonance:  
little gain from substrate Young's modulus

$$Y_{\text{Sapphire}} \sim 2.9 \times Y_{\text{Silicon}} \sim 2.2 \times Y_{\text{Silica}}$$

Granata 2013



- Substrate Brownian

$$S_{xx} \propto T \cdot \frac{1}{Y} \cdot \frac{1}{w f} \phi_s$$

- Substrate thermo-elastic

$$S_{xx} \propto T^2 \cdot \frac{\alpha^2 \cdot \kappa}{(C \cdot \rho)^2} \cdot \frac{1}{w^3 f^2}$$

- Coating Brownian

$$S_{xx} \propto T \cdot \frac{1}{Y} \cdot \frac{d}{w^2 f} \phi_c$$

DOMINANT  
TERM  
SO FAR

$$S_{xx} \propto T^2 \cdot \frac{1}{\sqrt{C_s \cdot \rho_s \cdot \kappa}} \cdot \frac{1}{w^2 \sqrt{f}} \cdot \left( \alpha_c \cdot d - \alpha_s \cdot d \frac{C_c \cdot \rho_c}{C_s \cdot \rho_s} - \beta_c \cdot \lambda \right)^2$$

- Coating thermo-optic

# Playing with numbers

- 1/30 from T
- 1/1.7<sup>2</sup> from  $w$
- 1/x from  $\varphi$

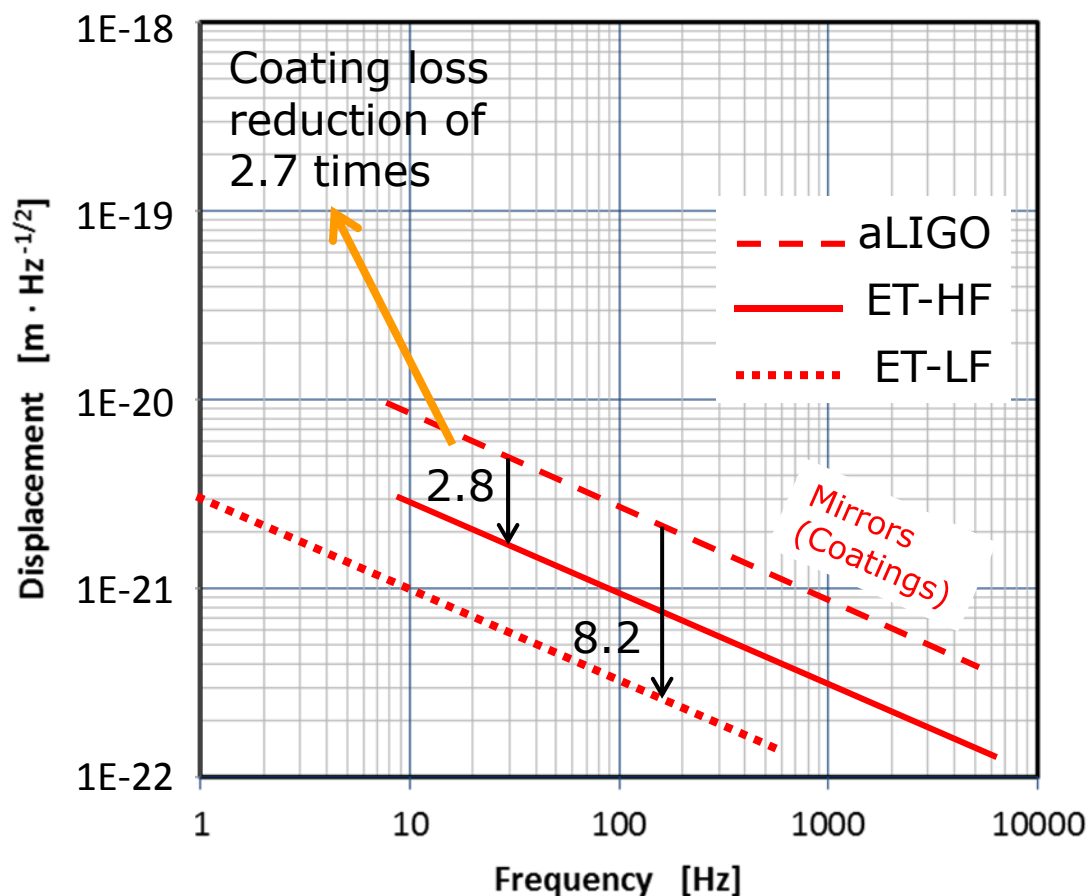
$\sqrt{S_{xx}}$  may decrease:

1.7 ·  $\sqrt{x}$  times @ 300 K

9.3 ·  $\sqrt{x}$  times @ 10 K

The 1.7 factor may increase further if the LG33 mode replaces the TEM00 in the cavities

Alberto Gatto's talk



# The factor x in the Amorphous Coatings

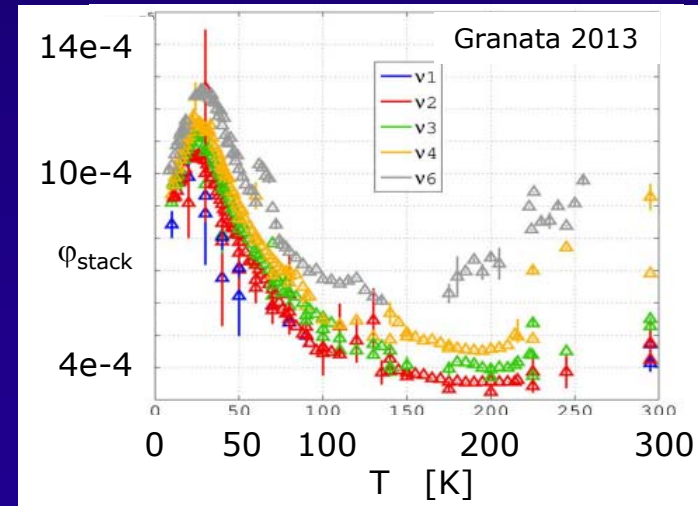
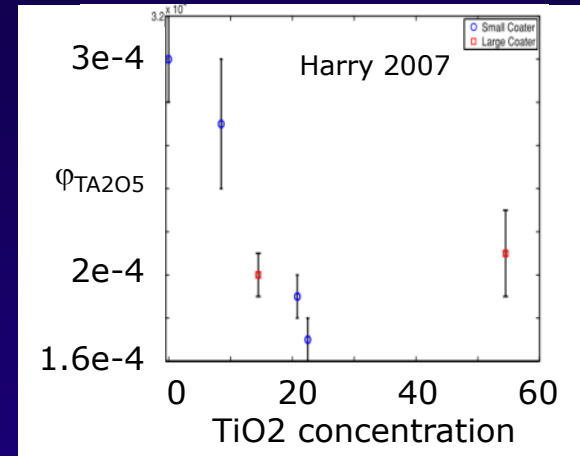
- Room temperature:

- Doping of Ta<sub>2</sub>O<sub>5</sub> with Ti developed at LMA
- Heat treatment
  - Effects at low T Martin 2010
  - At room T Massimo Granata's talk
- New materials and new doping
  - Steve Penn's research on ZrO<sub>2</sub>
- Nano-structured layers

Innocenzo Pinto's talk

- Low temperature

- $\phi_{\text{coat}}(10\text{K})$  is  $\sim 2 \times \phi_{\text{coat}}(300\text{K})$



# Structural studies

- To understand which are the relaxations responsible of thermal noise in amorphous materials

LIGO-Virgo meeting  
August 29<sup>th</sup> - Stanford

6<sup>th</sup> ET Symposium  
November 20<sup>th</sup> - Lyon

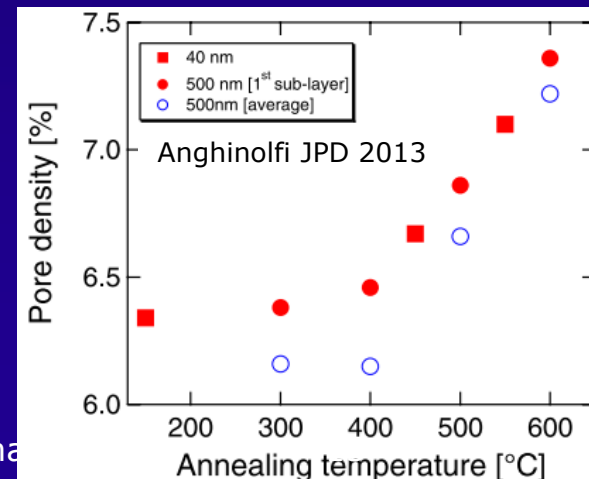
- The intelligent design of coatings follows

- Strategy

Riccardo Bassiri's talk

- Multi-technique analysis of static structure (x-rays and electronic diffraction) or dynamic (Raman)
- Correlation studies with mean properties like mechanical losses or refractive index
- Molecular dynamics models
- Controlled alteration of coatings

M. Granata's talk





# Crystalline coatings

- Low temperature operation

- AlGaP grown on Si
- AlGaAs grown on GaAs and lift off onto substrates

Iain Martin's talk

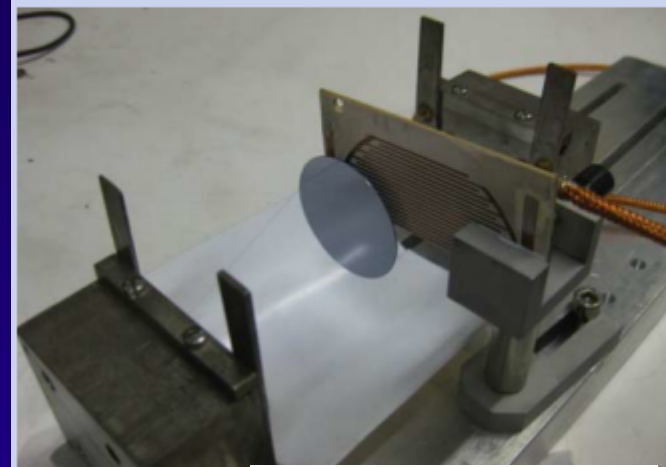
- Room temperature operation

- Lift off technique
- Thermo-optic noise seems under control

- Optical properties

- 5 to 13ppm absorp. measured on Cole's coatings
- Point like defects are sometimes present
- Scattering measurements

Coatings on Si disks: nodal support technique supported by 50  $\mu\text{m}$  wires



A. Lin's talk, GWADW 2013

Average coating loss (at 12K) calculated to be  $1.4 \times 10^{-5}$

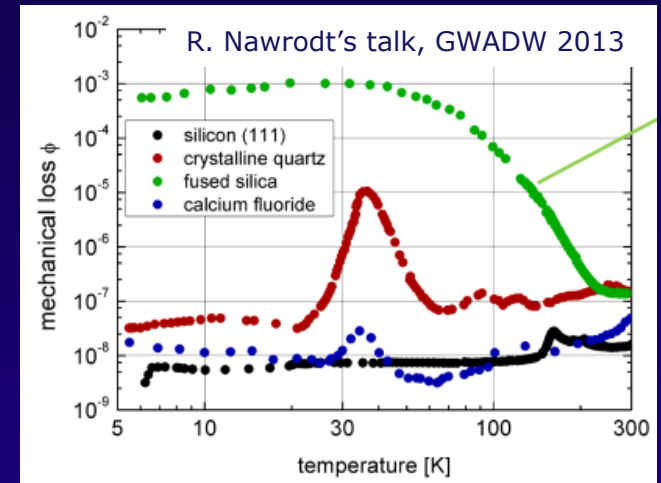
→ A factor of 45x lower than AdvLIGO  $\text{SiO}_2/\text{doped-Ta}_2\text{O}_5$  coating loss at 12K



# Crystalline substrates

- Silicon

- CZ vs FZ
- Homogeneity of impurities and of thermo-optical and mechanical parameters
- Effect of crystalline axis orientation



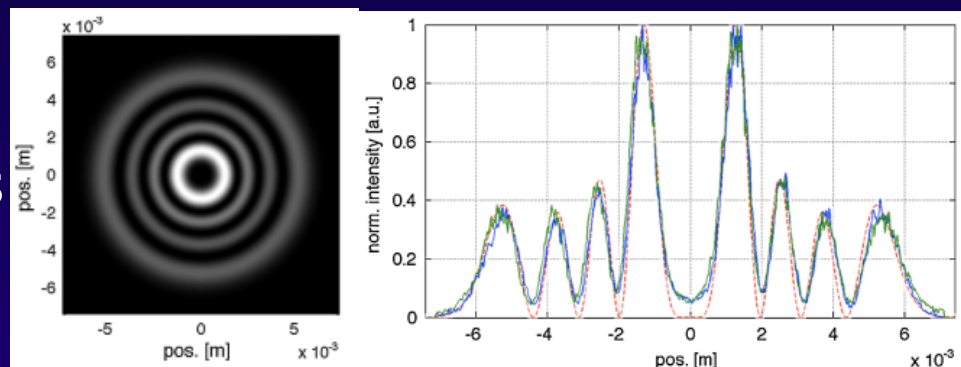
Gerd Hofmann's talk

- Sapphire

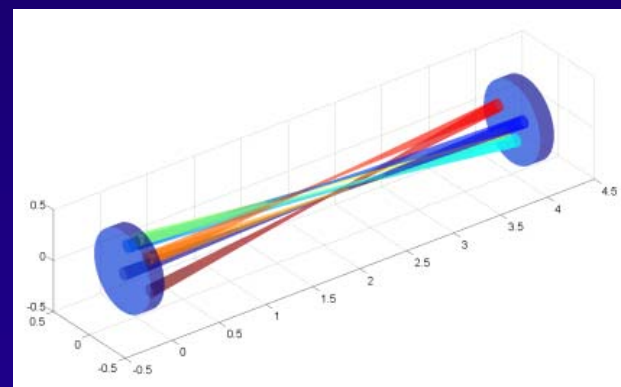
- Absorption

# Advancements in optics

- **LG33** Alberto Gatto's talk
  - Results from the latests experiments



- **Folded FP cavities** Stefan Balmer's talk
  - The history kicks back!
  - Sampling ideally the whole mirror surface
- **Centre of Mass detection**



# Parameters for different generations

	Virgo	AdV	LIGO	aLIGO	KAGRA	LIGO 3-R	ET-LF	ET-HF
Suspension material	Steel	Silica	Steel	Silica	Sapphire	Silica	Silicon	Silica
Suspension length, m	0.7	0.7	0.25	0.6	0.3	1.2	2	0.7
Wire diameter, $\mu\text{m}$	300	400	300	400	1600	566	3000	920
Bending p diameter, $\mu\text{m}$	300	800	300	800	1600	1624	?	1900
End mirror mass, kg	21	40	11	40	23	160	200	211
Input beam radius, mm	20	49	22	53	35	85	90	72
End beam radius, mm	55	58	22	62	35	100	90	72
Temperature, K	300	300	300	300	20	300	10	300

	Y [Mpa]	dE/E dT [1/K]	$\rho$ [g/cm <sup>3</sup> ]	$\alpha$ [1/K]	$\kappa$ [W/K m]	C [J/K kg]	n	dn/n dT [1/K]	$\phi$ structural
Steel	210	-1.32E-04	7.85	1.23E-05	58	480			1.80E-04
Fused silica	72	1.60E-04	2200	4.10E-07	1.38	716	1.45	8.00E-06	5.00E-09
Sapphire	300 K 392	-1.24E-04	3.99	5.49E-06	70	780			3.00E-09
	20 K 398	-1.41E-05		-2.91E-07	1000	0.71			
Silicon	300 K 130 / 188	-5.42E-05	2330		120				1.00E-08
	10 K				2200				3E-9 / 5E-9
Ta2O5 doped	300 K 140		und. 7390	3.60E-06	33	280	2.06	1.40E-05	2.44 E-4
	10 K								
SiO2	300 K 70		2470	5.00E-07	1.38	745	1.45	8.00E-06	4.60E-05
	10 K								
GaAs	300 K 85.5 / 140		5.32	5.80E-06	51	327	3.3	4.50E-05	Mirror 2.5E-5
	10 K			-3.84E-09	2000	14.6			
GaP	300 K 103		6.15	4.73E-06	110	444	3.05		
	10 K								Mirror 1.4E-5

# Conclusion

- Progresses are constants
- There are good chances to reach the expected noise levels for future detectors
- A POSITIVE COMPETITION BETWEEN MATERIAL RESEARCH AND OPTICS IS NECESSARY

