

MIST in Quantumland

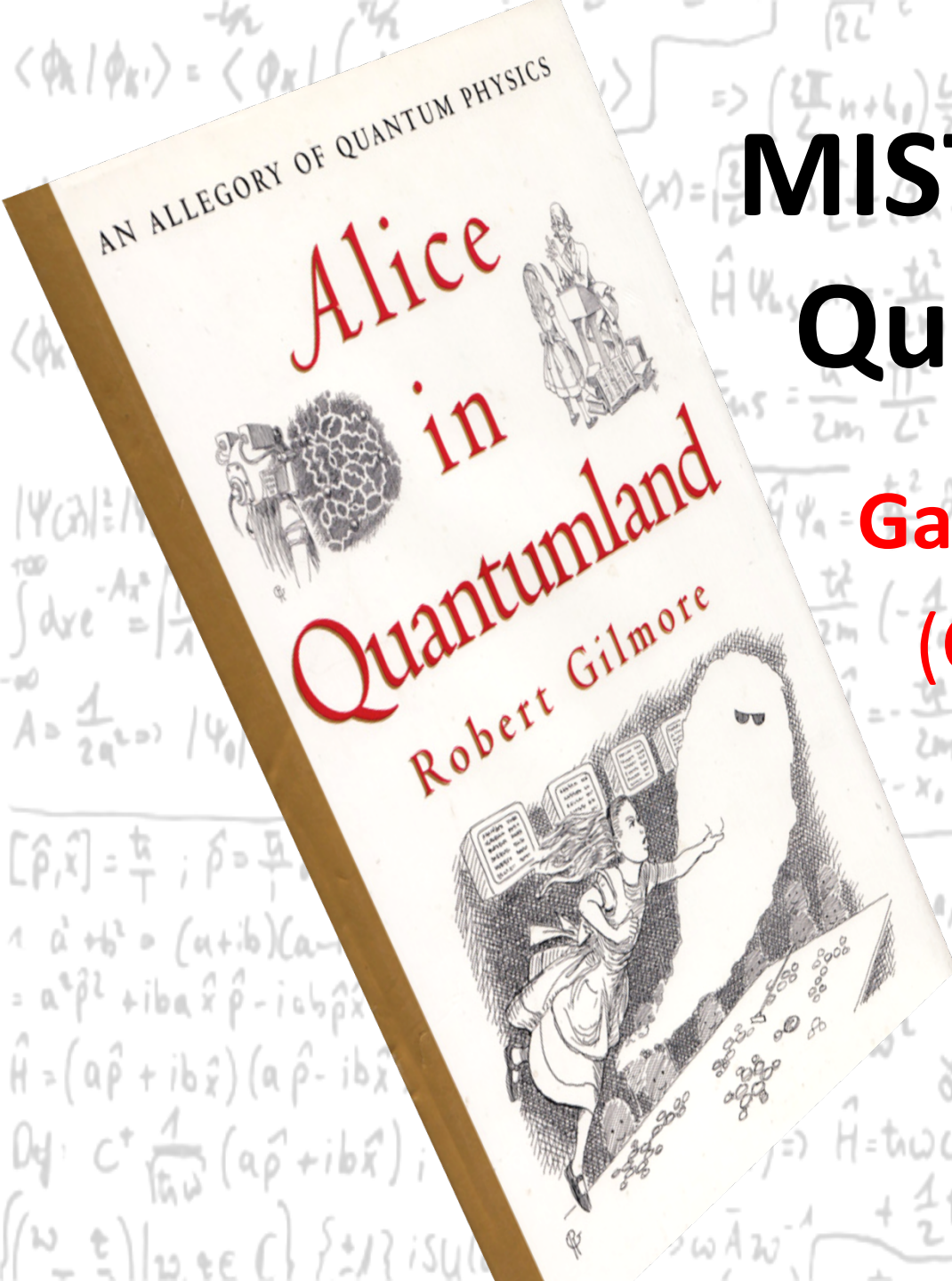
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GWADW 2014

Takayama

May 28th 2014

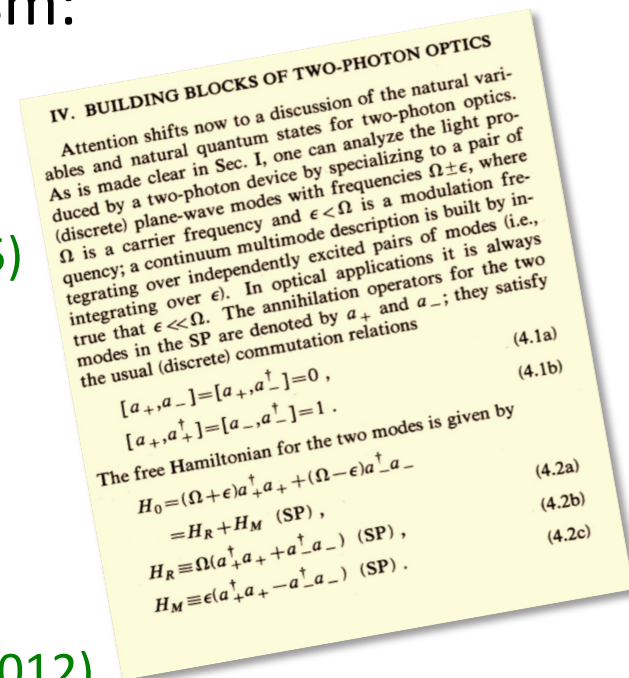
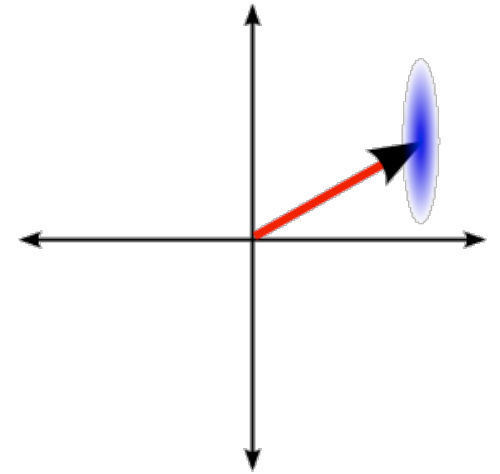
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- This is not a “**What is MIST?**” talk...
 - LVC March 2014 (LIGO-G1400322)
 - LIGO seminar (LIGO-G1400061)
 - MIST web page
<http://sourceforge.net/projects/optics-mist/>
 - G. Vajente, *Class. Quantum Grav.* **30**, 075014 (2013)
 - G. Vajente, *Class. Quantum Grav.* **31**, 075005 (2014)
- Radiation pressure is simulated in MIST since September 2013
- In the last version, **quantum noise** computation is also included
- This talk is about **how** quantum noise is computed in MIST



- We often explain quantum noise in the framework of squeezed light with the “ball-on-the-stick” description
- Mathematically, quantum noise is often described in the “two-photons” formalism:
 - Seminal papers
 - Caves et al. Phys. Rev. A 31, 3068 (1985)
 - Schumaker et al. Phys. Rev. A 31, 3093 (1985)
 - A more modern approach
 - Kimble et al. Phys. Rev. D 65, 022002 (2001)
 - Corbitt et al. Phys. Rev. A 72, 013818 (2005)
 - A good (and long) review
 - Danilishin et al. Living Rev. Relativity 15, 5 (2012)



- MIST and most simulations treat audio frequency motions in the “audio sidebands” formalism

$$E(t) = e^{i\omega_0 t} (E_0 + E_+ e^{i\Omega t} + E_- e^{-i\Omega t})$$

where ω_0 is the laser carrier frequency and Ω is the signal frequency (for example motion of a mirror)

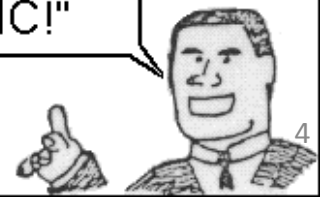
- In quantum noise papers you see it written in a slightly more complex way:

$$E(t) = \sqrt{\frac{\hbar\omega_0}{Ac}} e^{i\omega_0 t} \int_0^\infty \frac{d\Omega}{2\pi} [a_+(\Omega)e^{i\Omega t} + a_-(\Omega)e^{-i\Omega t}]$$

- But as long as you ignore that the a ’s are quantum operators, everything is fine...

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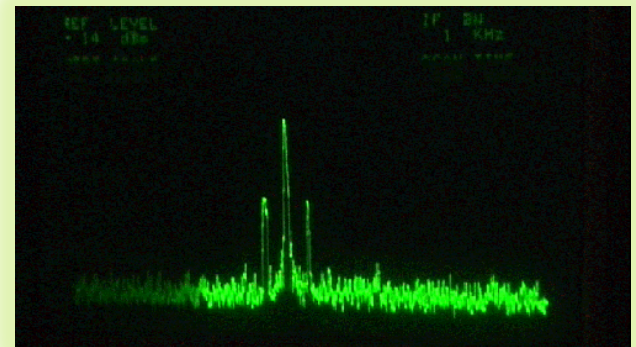
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- The two-photons formalism mixes the audio sidebands, to take into account phenomena where correlated pairs of photons are created (like OPAs in squeezers)

$$a_1 = \frac{a_+ + a_-^+}{\sqrt{2}} \quad a_2 = \frac{a_+ - a_-^+}{\sqrt{2}}$$

- This is just a linear combination of the coefficients we normally use in the audio sidebands approach



$$\begin{pmatrix} E_+ \\ E_-^+ \end{pmatrix} = \sqrt{\frac{h\omega_0}{16\pi}} \begin{bmatrix} 1 & i \\ 1 & -i \end{bmatrix} \begin{pmatrix} a_+ \\ a_- \end{pmatrix} = M \begin{pmatrix} a_+ \\ a_- \end{pmatrix}$$

- Since our IFOs are linear systems for the audio sidebands, we can compute their input-output relations using either formalism, and convert back to the other one

Quantum noise – basic computation

Quantum vacuum fields enter from all open ports of the IFO

The IFO acts as a linear operator and can mix the two quadratures

And we get some vacuum fields at the output, impinging on our PD

The signal at the photodiode output (at audio frequency) is a linear combination of the two quadratures

$$\begin{pmatrix} a_1 \\ a_2 \end{pmatrix} \xrightarrow{\text{orange wavy arrow}} T = \begin{bmatrix} \cdot & \cdot \\ \cdot & \cdot \end{bmatrix} \xrightarrow{\text{green curly arrow}} \begin{pmatrix} b_1 \\ b_2 \end{pmatrix} \xrightarrow{\text{blue wavy arrow}} S = C_1 b_1 + C_2 b_2$$

To get the power spectral density of the noise on the signal, we have to “average” the quantum operators on the input vacuum field

$$S(\Omega)\delta(\Omega - \Omega') = \langle in | S(\Omega)S(\Omega')^+ | in \rangle$$

$$S(\Omega) = [C_1 \ C_2] T S_v T^+ \begin{bmatrix} C_1^* \\ C_2^* \end{bmatrix}$$

S_v is the vacuum cross-spectral density matrix (which is unity for unsqueezed vacuum)



- MIST **does not use the two photons formalisms**
- Quantum vacuum fields can be injected anywhere
 - They can be converted into the corresponding audio sideband components using the “base change” matrix M
- Then MIST can propagate as usual the audio sidebands to all output ports, including radiation pressure effects

$$\begin{pmatrix} F_+ \\ F_-^+ \end{pmatrix} = T \begin{pmatrix} E_+ \\ E_-^+ \end{pmatrix}$$

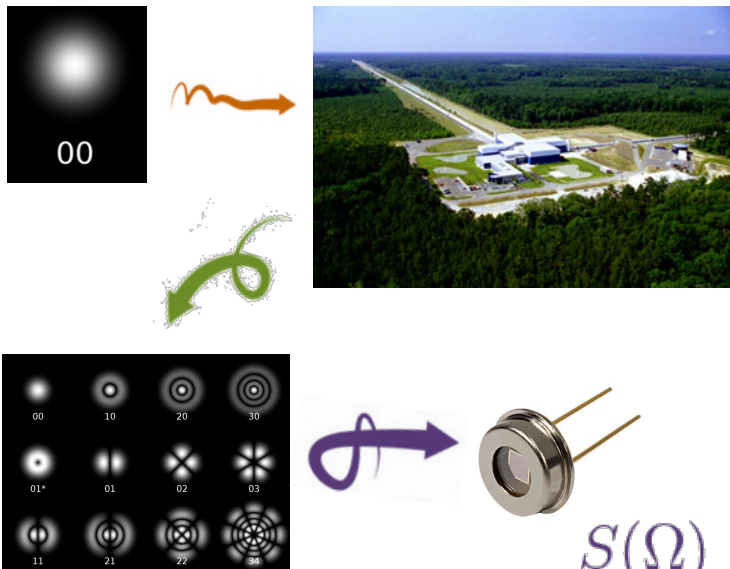
- The (audio frequency) signal on a photodetector is a linear combination of the audio sidebands

$$S(\Omega) = C_+ F_+ + F_-^+ C_-$$

where in general the fields F are vectors in a HOM basis



- Each **higher order mode** in the input vacuum field is assumed to be **independent** of the others
- It must be propagated **separately**:
 - MIST computes the output of the IFO for each one of the HOMs in input separately, including all possible couplings among modes
- Each HOM in input contributes to quantum noise in output
 - All contributions are summed **incoherently**

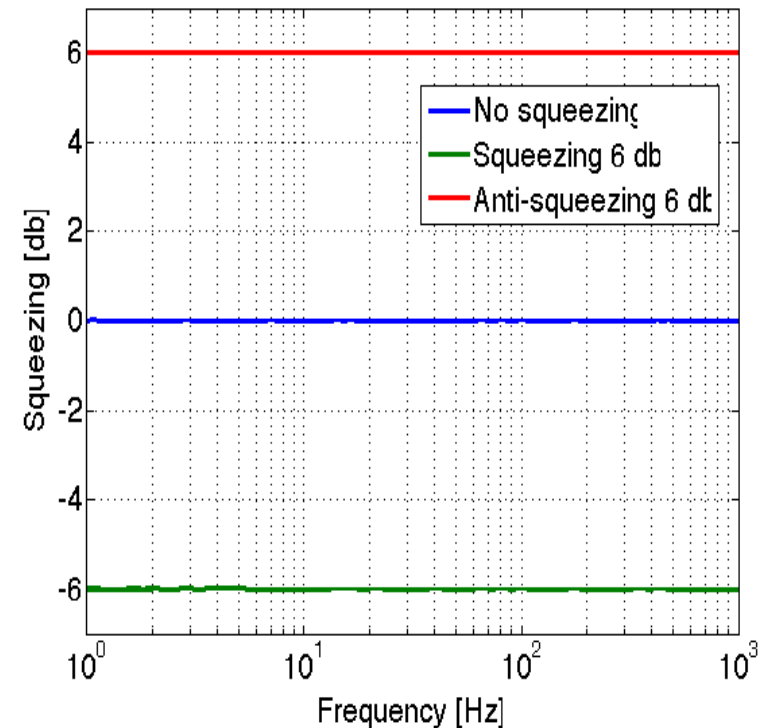


$$S(\Omega) = \begin{bmatrix} (C_+ & C_-^T) T_{00 \rightarrow HOM_s} M \end{bmatrix} \cdot S_v \cdot \begin{bmatrix} M^+ T_{00 \rightarrow HOM_s}^+ \begin{pmatrix} C_+^+ \\ C_-^* \end{pmatrix} \end{bmatrix}$$

One contribution from each HOM of the input vacuum field

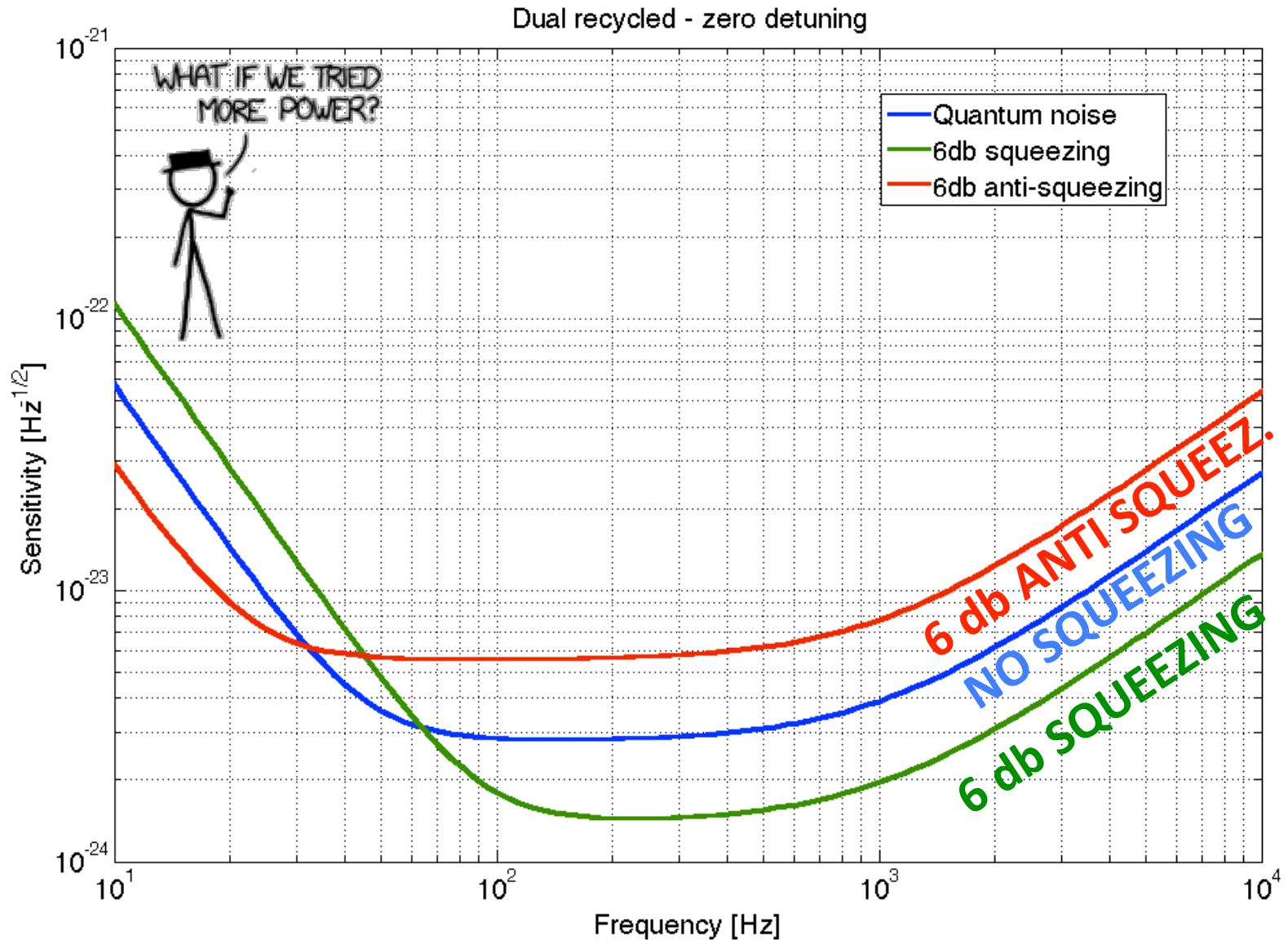
$$S(\Omega) = C_+ F_+ + F_-^+ C_-$$

- Presently, the user must decide at which ports quantum vacuum fields must be injected
 - MIST does not include automatic injection of vacuum at all lossy ports
 - In the future this will be an option, together with automatic computations of losses due to clipping
- Arbitrary squeezing factor and angle can be injected in any point inside the optical system

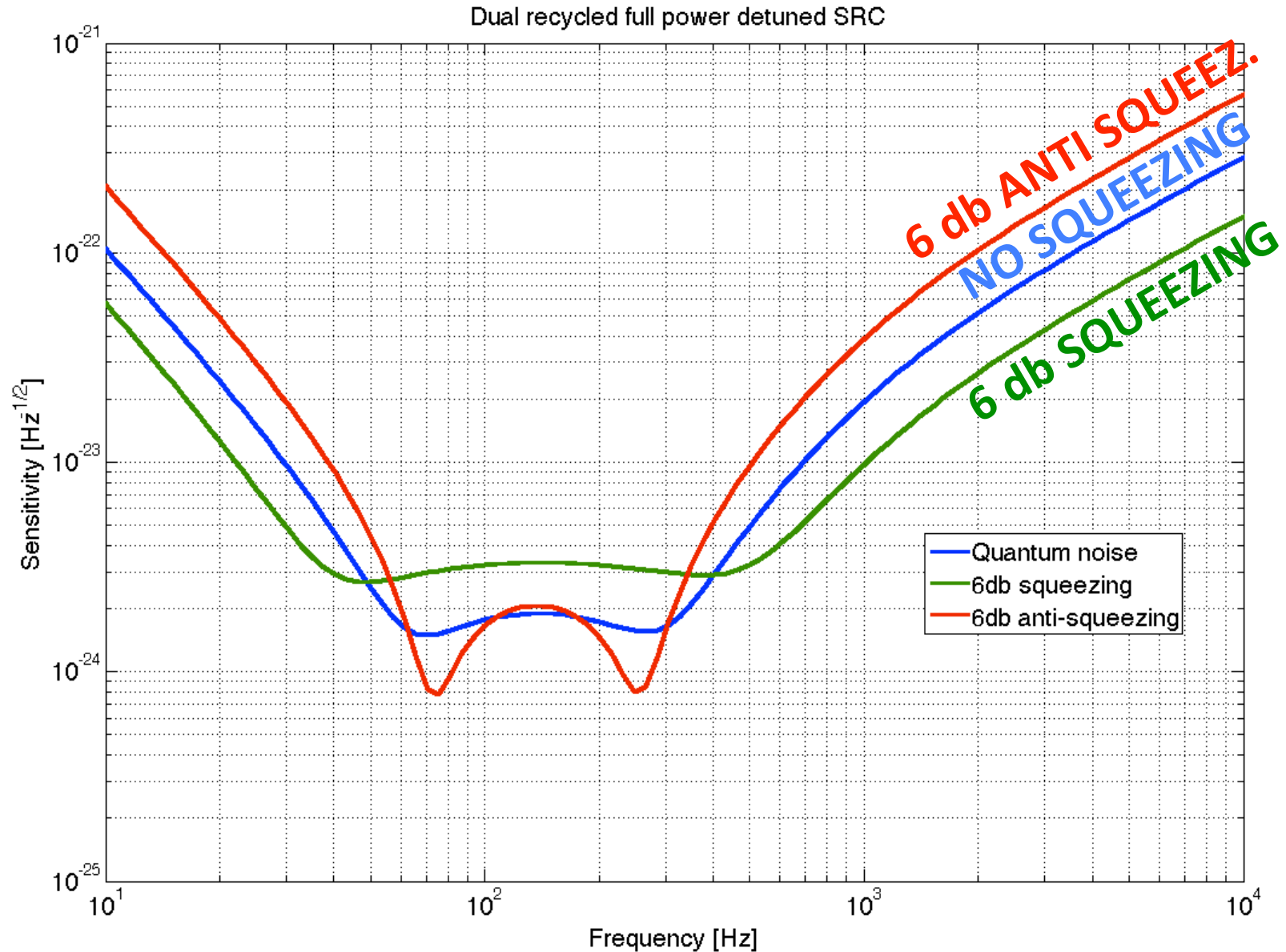


```
vacuum name node
  db=squeezing-factor-db
  angle=rotation-angle-deg
  w=vacuum-mode-beam-size
  R=vacuum-mode-curvature
  ampl=amplitude-relative
  from=object-to-set-direction
  into=object-to-set-direction
```

Dual recycling with squeezing /1

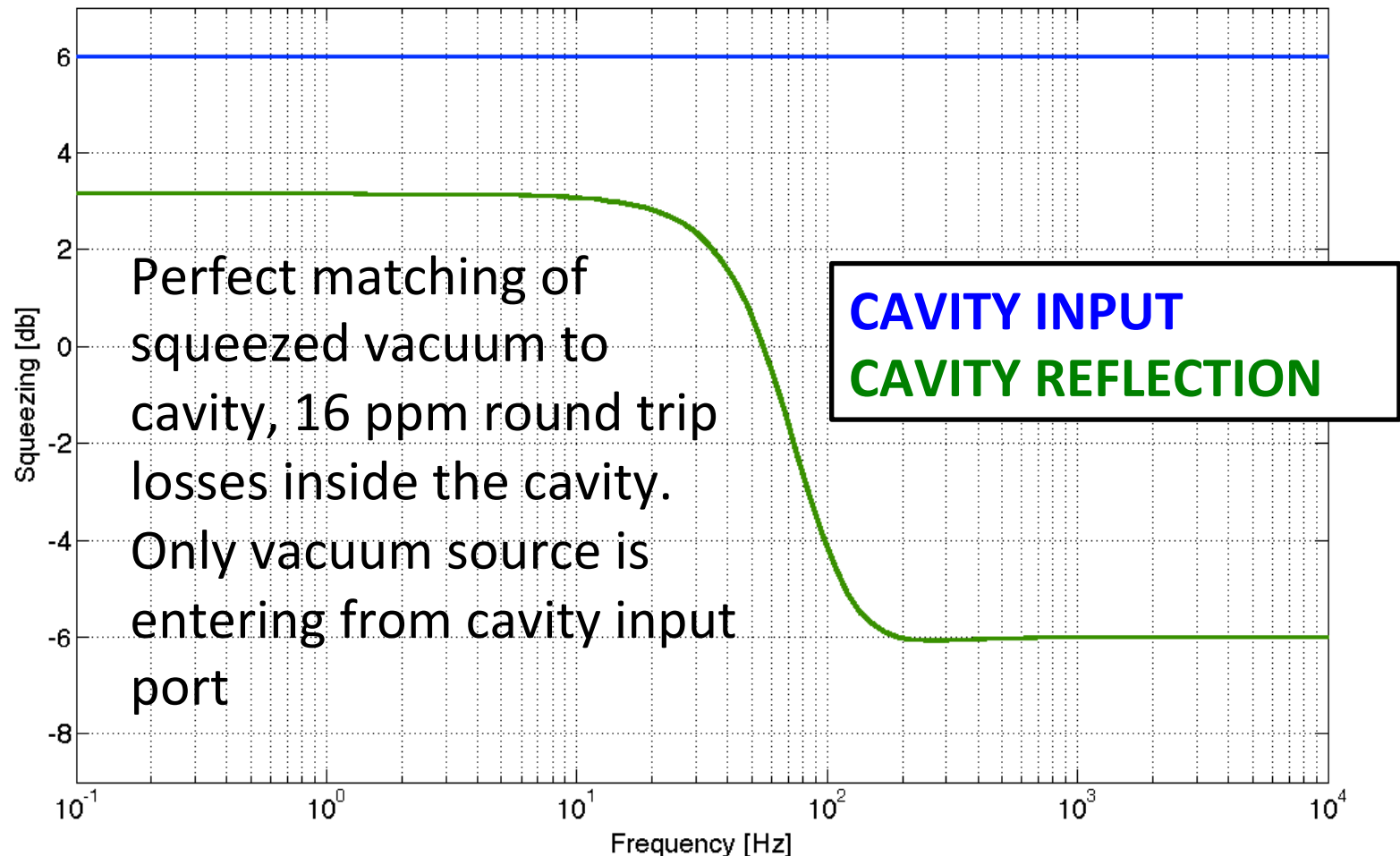


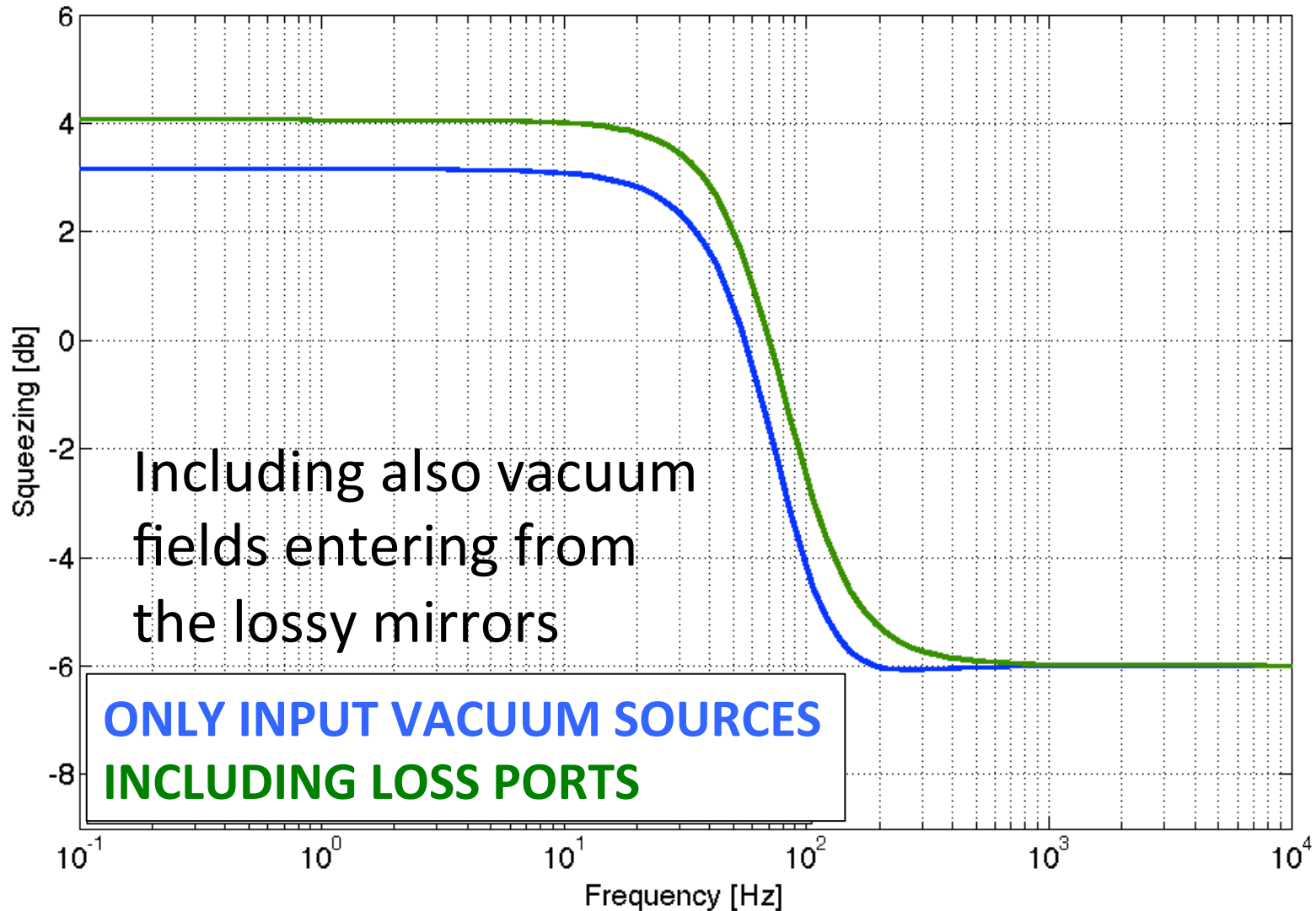
Dual recycling with squeezing /2

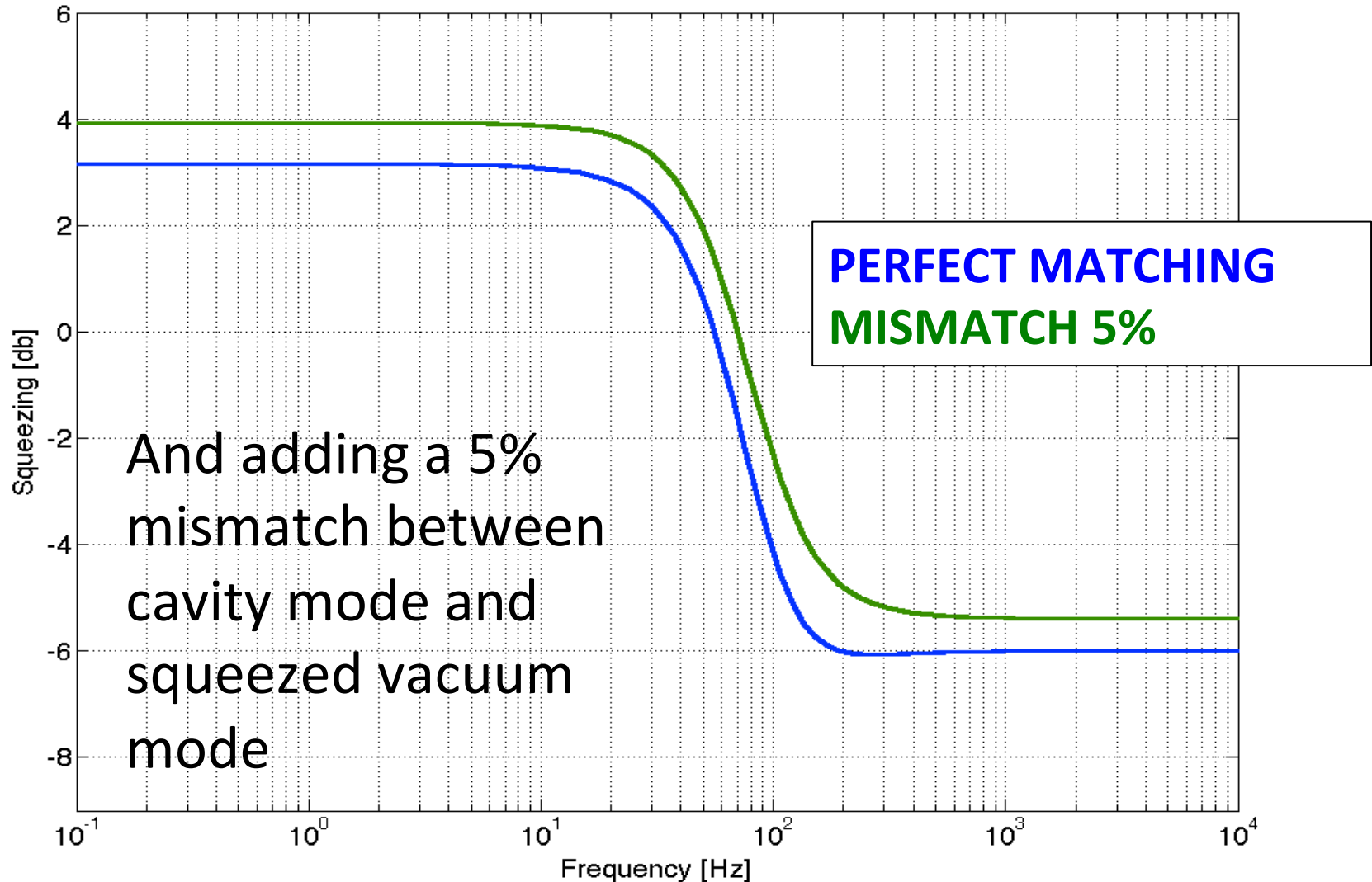


A filter cavity /1

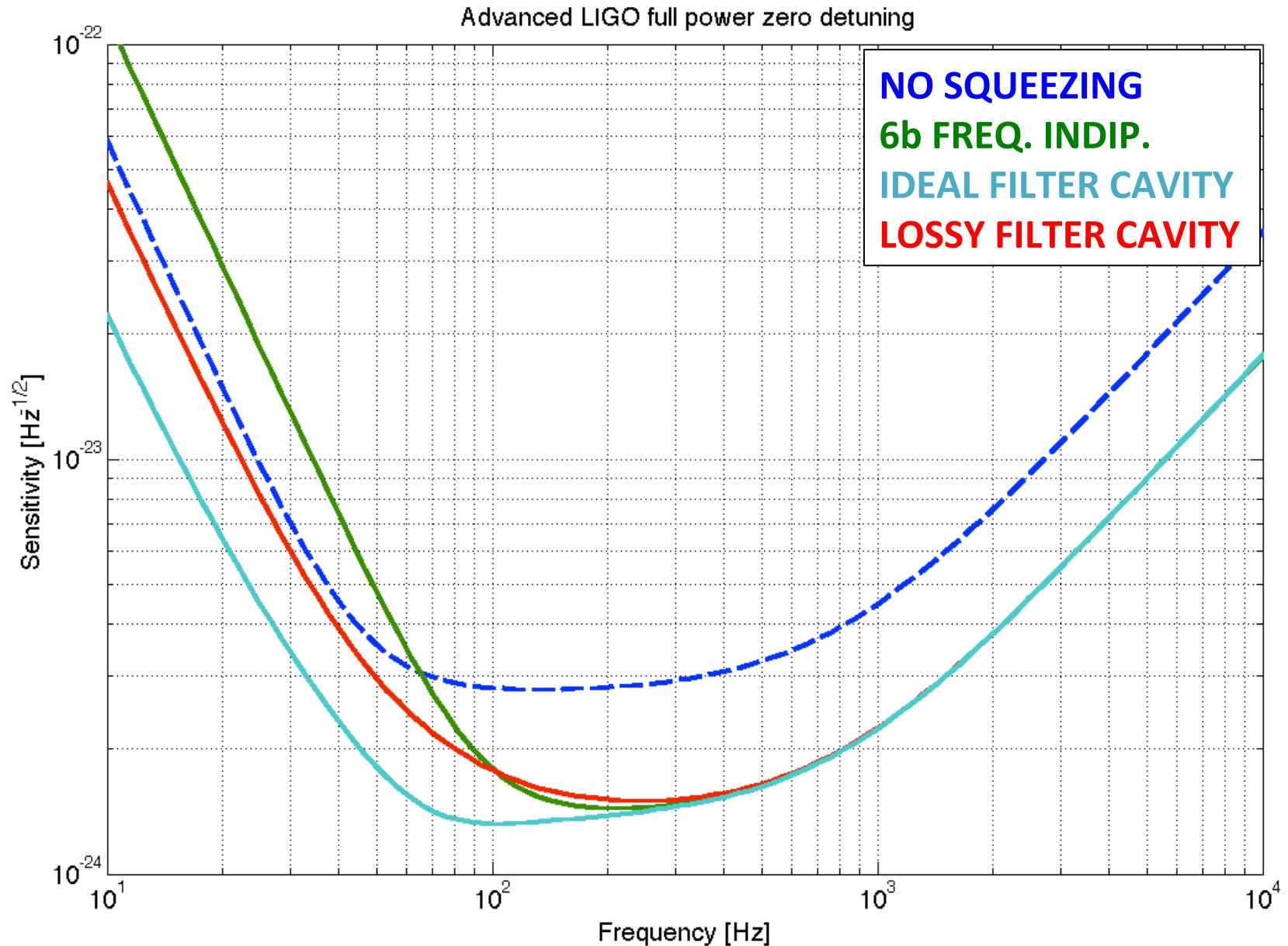
- Parameters from *Evans et al. Phys. Rev. D 88, 022002 (2013)*



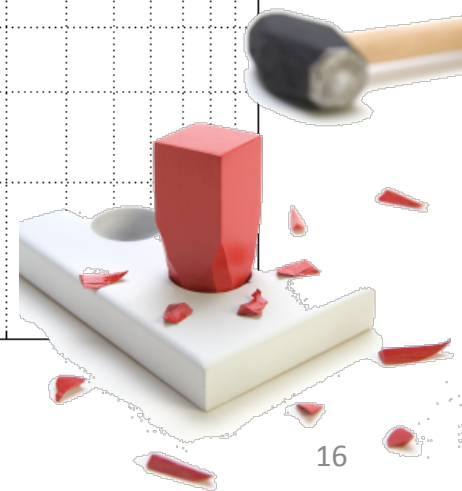
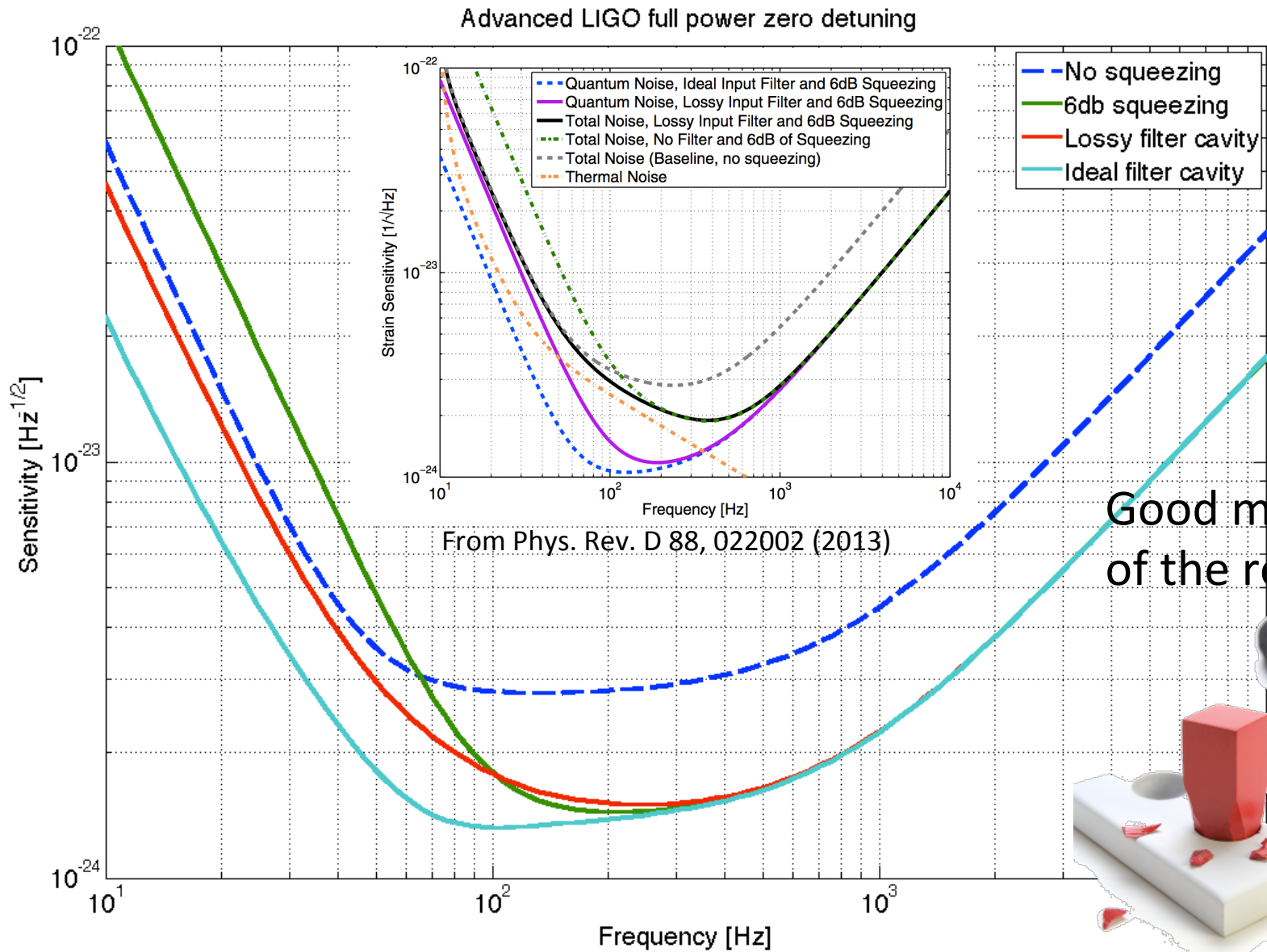




Advanced LIGO with filter cavity



Advanced LIGO with filter cavity



MIST entered in the quantum land, who knows what's there?

If you're interested in simulating quantum noise with MIST, contact me (vajente@caltech.edu) and download MIST:

<http://sourceforge.net/projects/optics-mist/> *

and subscribe to the **google group**:

<https://groups.google.com/forum/#!forum/mist-users-group>



TO DO SOON:

- add automatic insertion of vacuum fields at all lossy ports (including clipping points)

* (almost) all the examples shown in this talk are included in the MIST package documentation