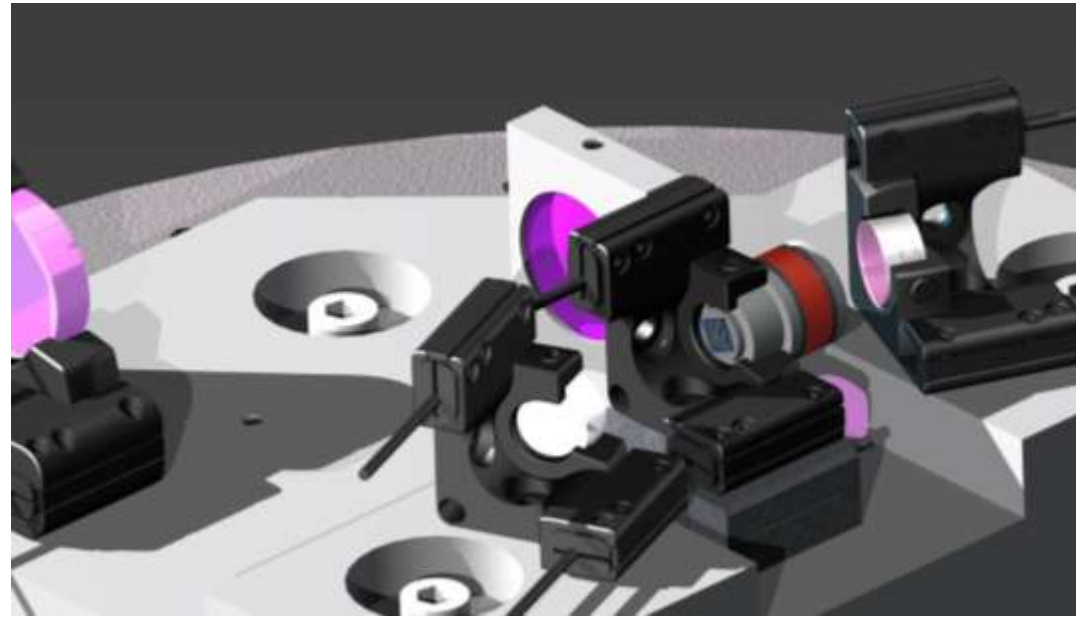


Dissipative Optomechanical Coupling



Andreas Sawadsky, H. Kaufer,
R. Moghadas Nia, and R. Schnabel

Max Planck Institute for Gravitational Physics,
Leibniz Universität Hannover

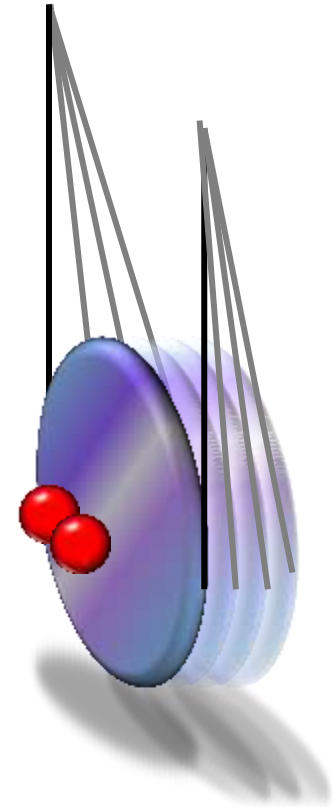
Theoretical work:
S. Tarabrin and K. Hammerer

Max Planck Institute for Gravitational Physics,
Institute for Theoretical Physics, Leibniz Universität
Hannover



Optomechanical coupling: Back-action

- Radiation pressure of light field leads to a force on a test mass (mirror)
- Quantum back-action $\rightarrow$ RPN+SN $\rightarrow$ SQL





Dynamical back-action



Effective FP-Cavity

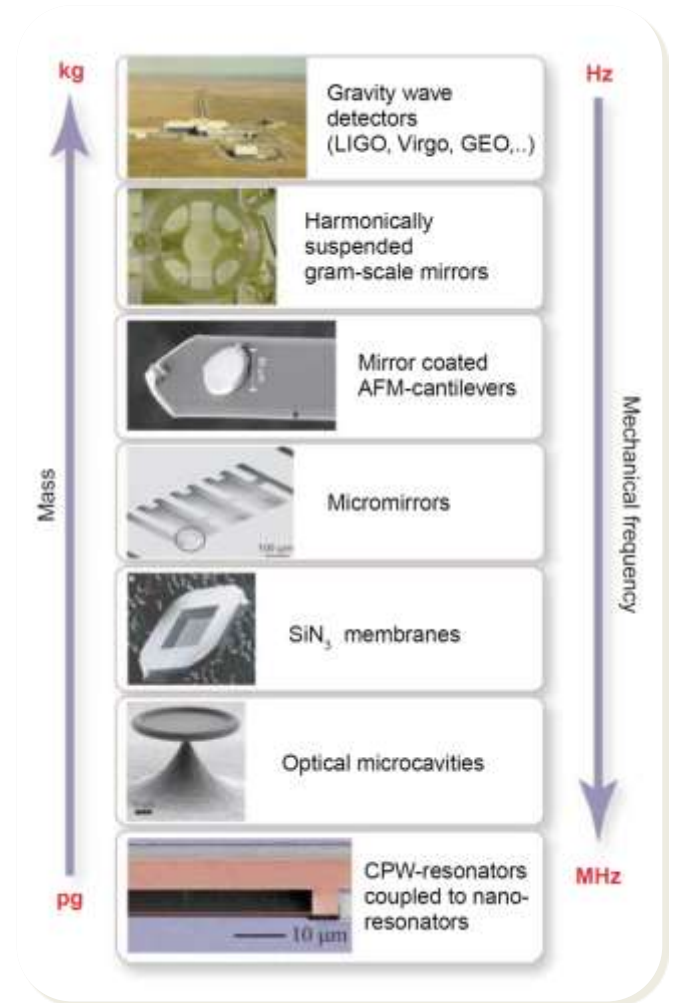
- Micro-and nano oscillators
- Moveable mirror

[S. Gigan et al., *Nature* 444, 67 (2006), T. Corbitt et al., *PRL* 98, 150802 (2007), J. D. Thompson et al., *Nature* 452, 72 (2008), etc...]



Interferometer (GWD)

- Scaling law: noise + dynamics operated on dark port are equivalent to those of a FP-Cavity
[A. Buonanno, Y. Chen, *Phys. Rev. D* 67, 062002 (2003)]



T. J. Kippenberg, K. J. Vahala, *Science* 321, 1172 (2008)

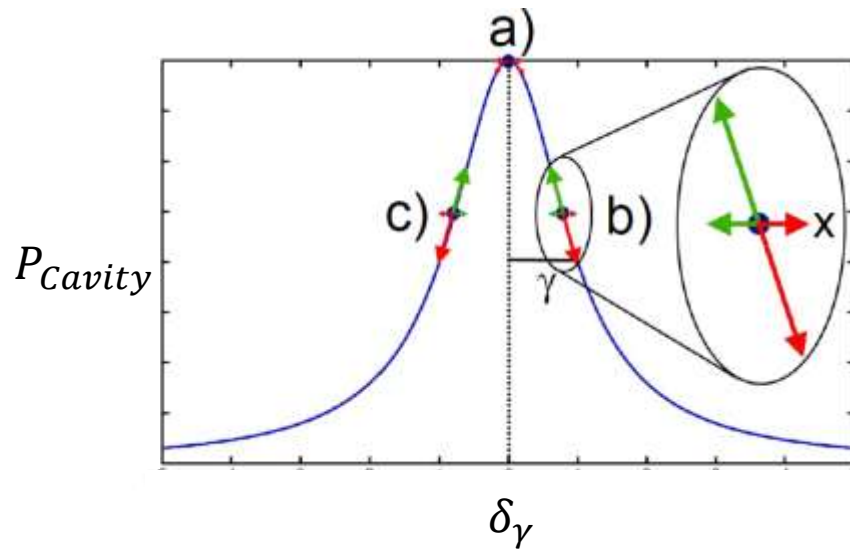
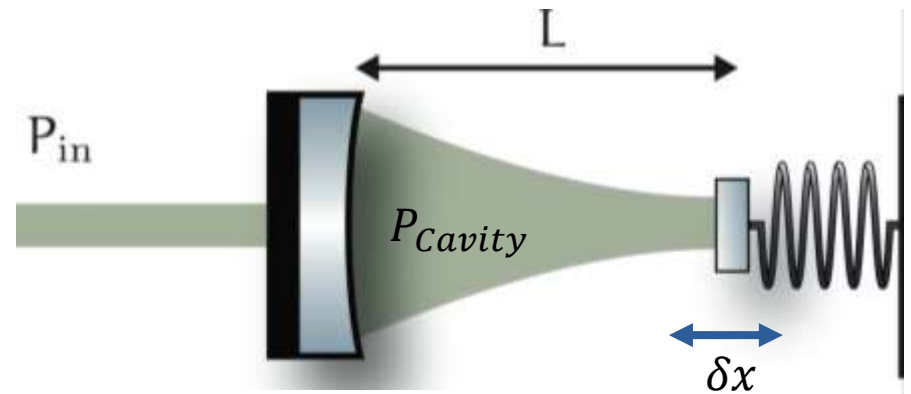


Dynamical Back-Action: Optical Spring

Interaction Hamiltonian: $\mathcal{H}_{int} = g_{\omega} \hat{x} \hat{a}^+ \hat{a}$

dispersive coupling: $g_{\omega} = \frac{\delta\omega}{\delta x}$

Resonance condition: $L = \frac{n\pi c}{\omega_L}$



a) $\delta = 0$:
 P_{cavity} depends only in
 2. order of magnitude to δx .

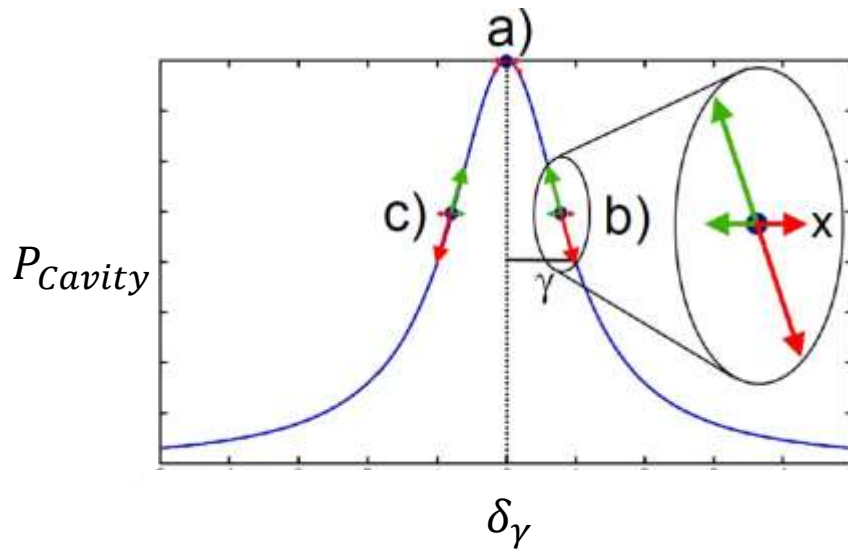
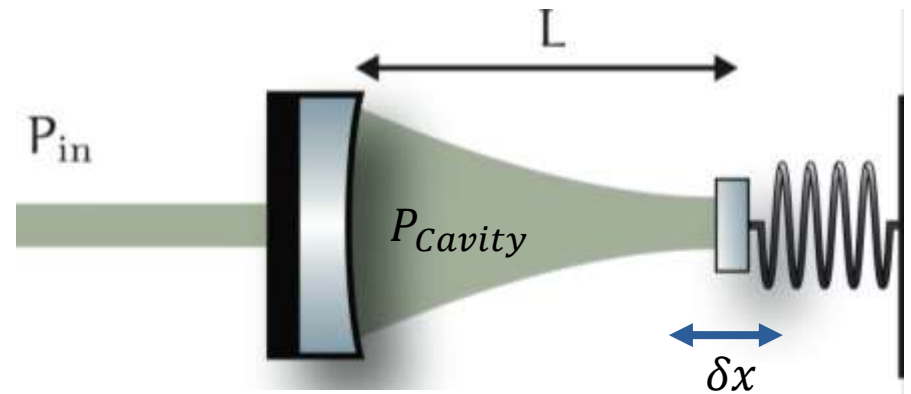


Dispersive Coupling

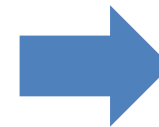
Interaction Hamiltonian: $\mathcal{H}_{int} = g_{\omega} \hat{x} \hat{a}^+ \hat{a}$

dispersive coupling: $g_{\omega} = \frac{\delta\omega}{\delta x}$

Resonance condition: $L = \frac{n\pi c}{\omega_L}$



b/c) $\delta \neq 0$:
linear dependence
 $P_{cavity} \propto \delta x \Rightarrow F_{RP} \propto \delta x$

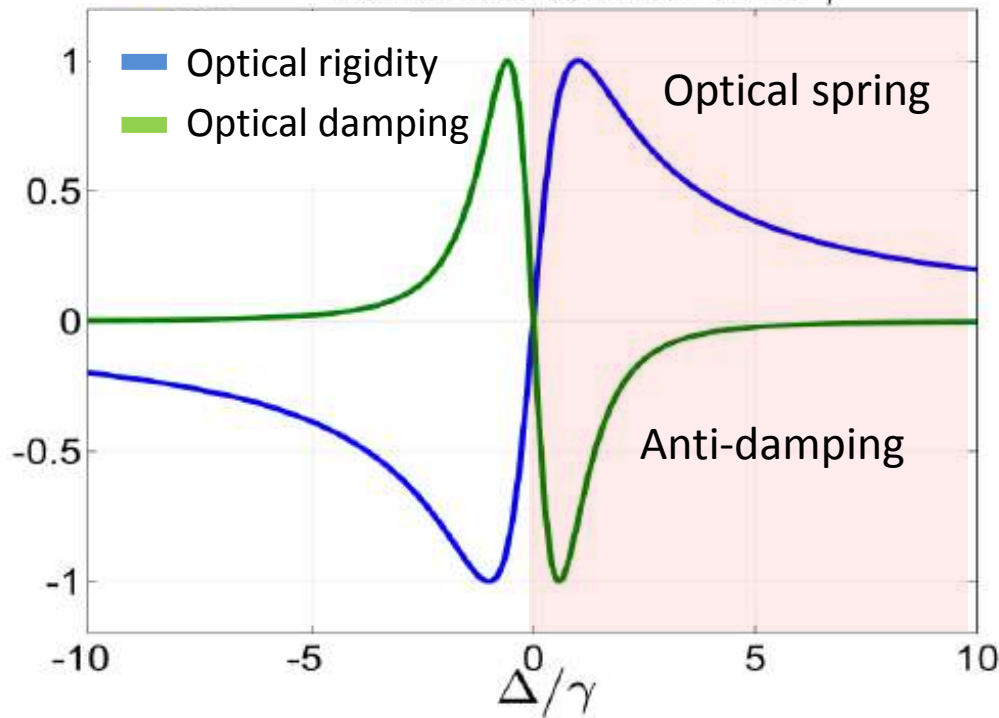


Optical spring



Dynamical Back-Action in GWD

GWD scenario: $\Omega < \gamma$



$$F(\Omega) = -\delta x [K(\Omega) - 2i\Omega\Gamma(\Omega)]$$

$K(\Omega)$: *optical rigidity*

$\Gamma(\Omega)$: *optical damping*



Optical Spring in GWD is unstable!

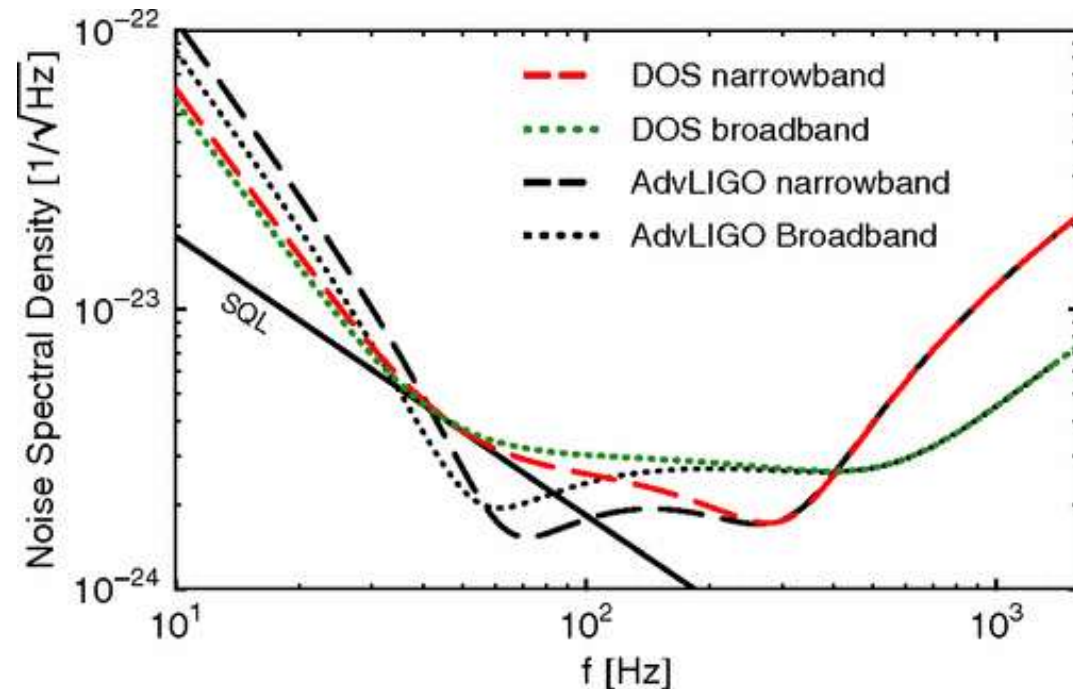
V.B. Braginsky et al., *Phys. Lett. A* **232**, 340 (1997);

F.Ya. Khalili, *Phys. Lett. A* **288**, 251 (2001);

A. Buonanno, Y. Chen, *Phys. Rev. D* **65**, 042001 (2002)



Dynamical Back-Action in GWD



Optical Spring in GWD is unstable!

- control/feedback needed for stabilization
- 2 Laser drives can create stable spring and sensitivity improvement

H. Rehbein et al., PRD 78, 062003 (2008)

T. Corbitt et al., PRL 98, 150802 (2007)

H. Rehbein et al., PRD 78, 062003 (2008)



Dynamical Back-Action in GWD



Dynamical back-action on dark-port

- 2nd generation of GWD will operated slightly off dark-port

f [Hz]

H. Rehbein et al., PRD 78, 062003 (2008)



What about dynamic back-action off dark-port?



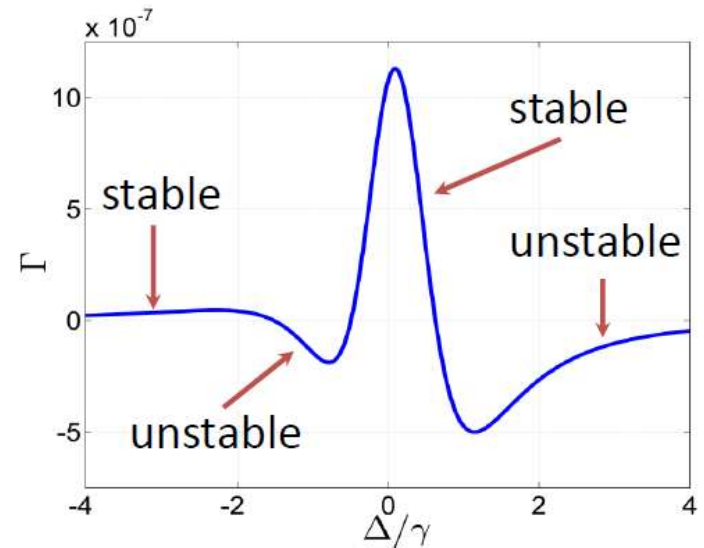
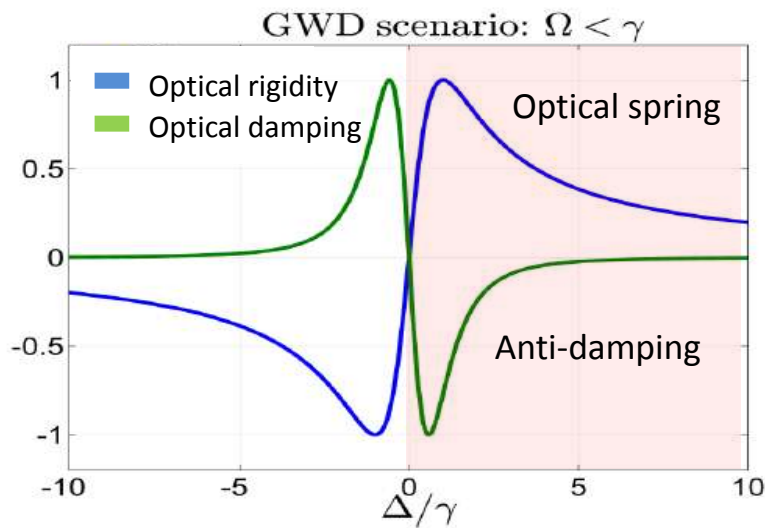
Additional instabilities

*S. Tarabrin et al.,
Phys. Rev. A 88, 023809 (2013)*

On dark-port



Far away
Off dark-port



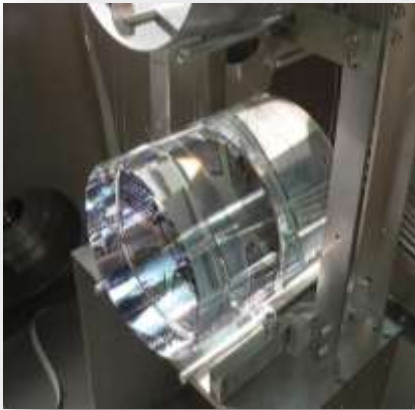
Not crucial for GWD!



LARGE SCALE to small scale experiment

Gravitational Wave Detector (GWD)

~ 200 W high input power
~ 40 Kg free test Masses
~ 100% power reflectivity
~ 1 Hz resonant frequency



Suspended fused silica mirror

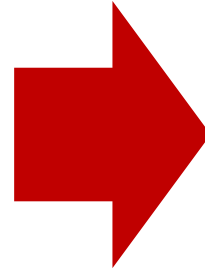
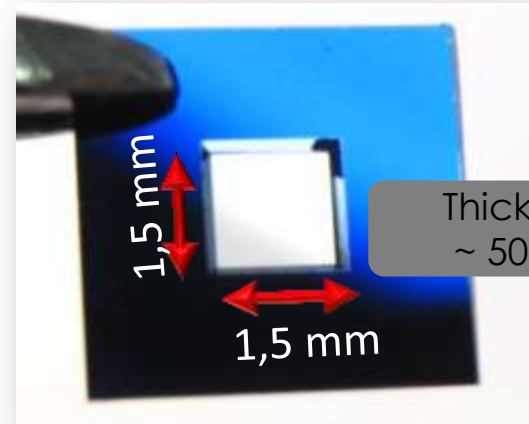


Table top experiment

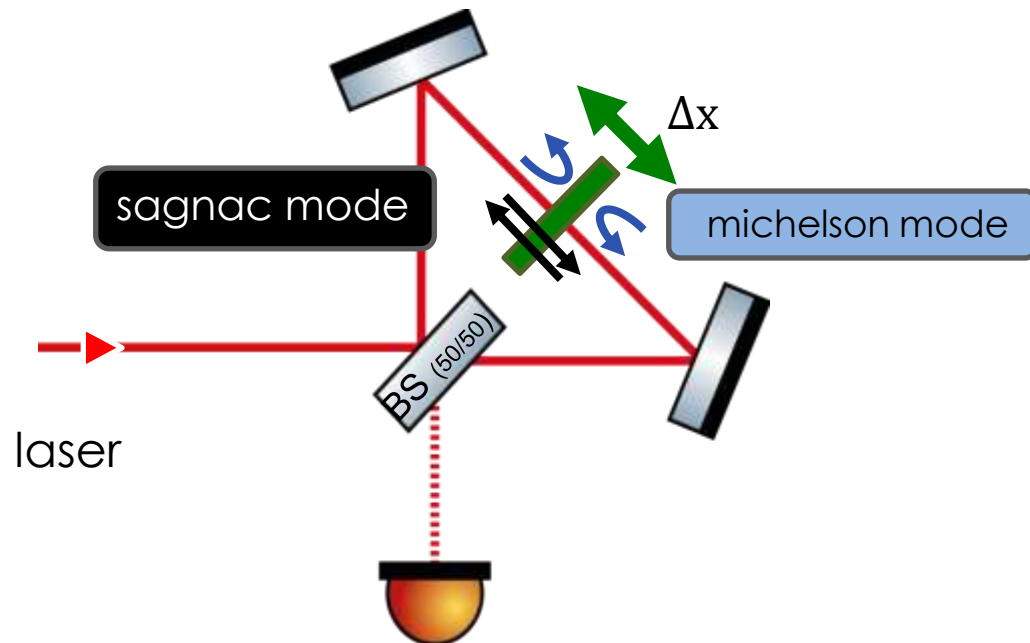
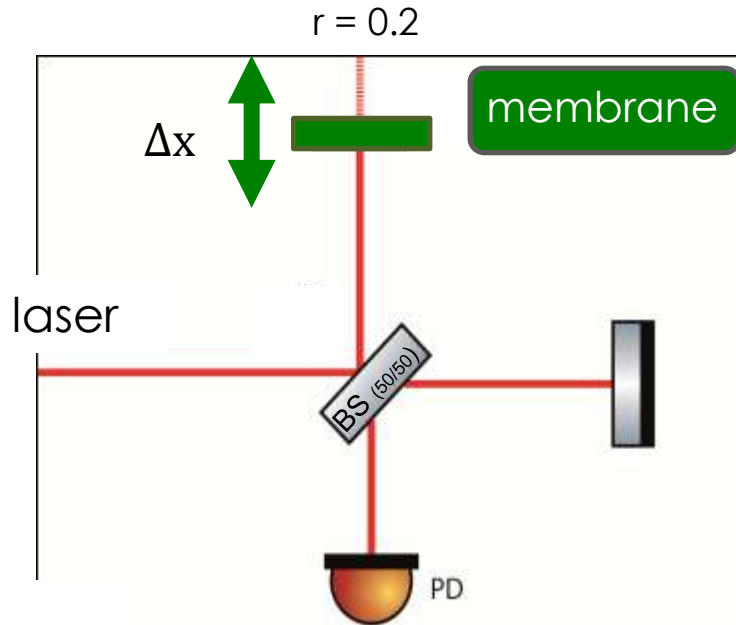
~ 1 mW-1 W input power
~ 100 ng test mass
~ 20 % power reflectivity
~ 100 kHz resonant frequency



SiN – Silicon Nitride membrane



Michelson – Sagnac Interferometer

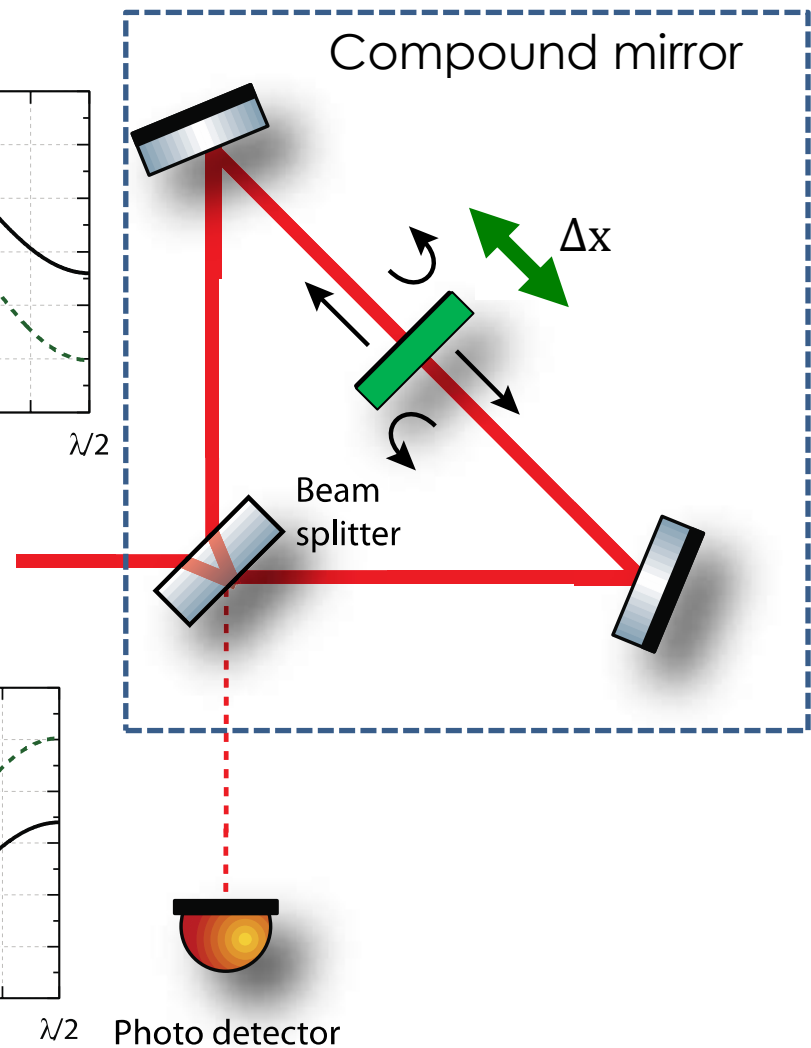
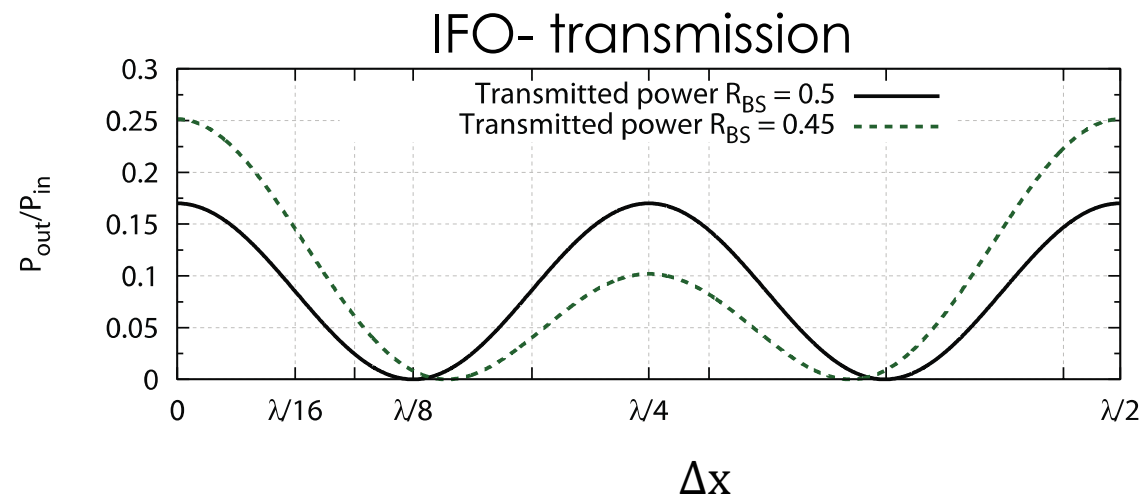
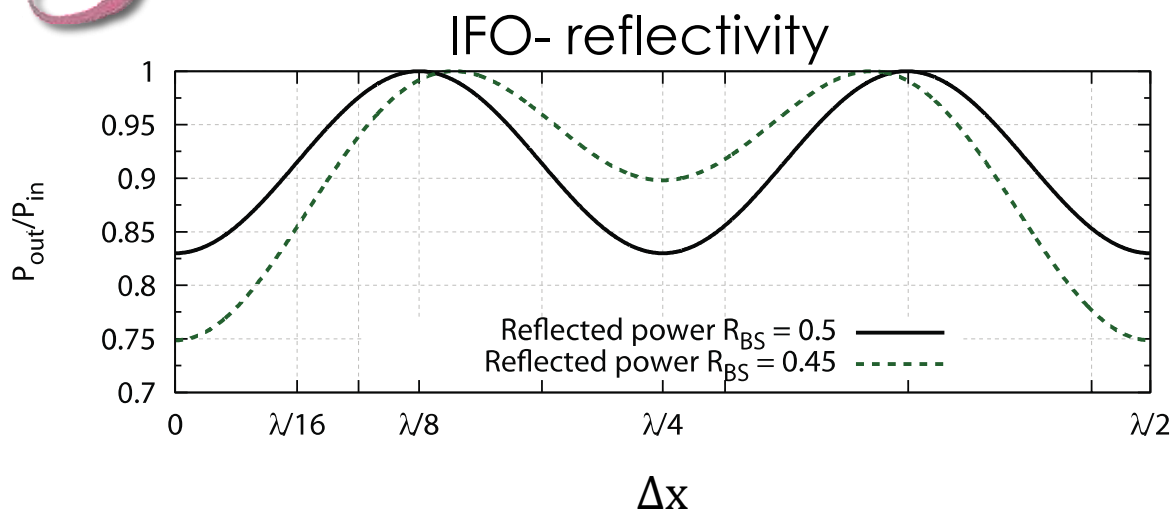


K. Yamamoto et al., Phys. Rev. A **81**, 033849 (2010)

D. Friedrich et al., New J. Phys. **13**, 093017 (2011)

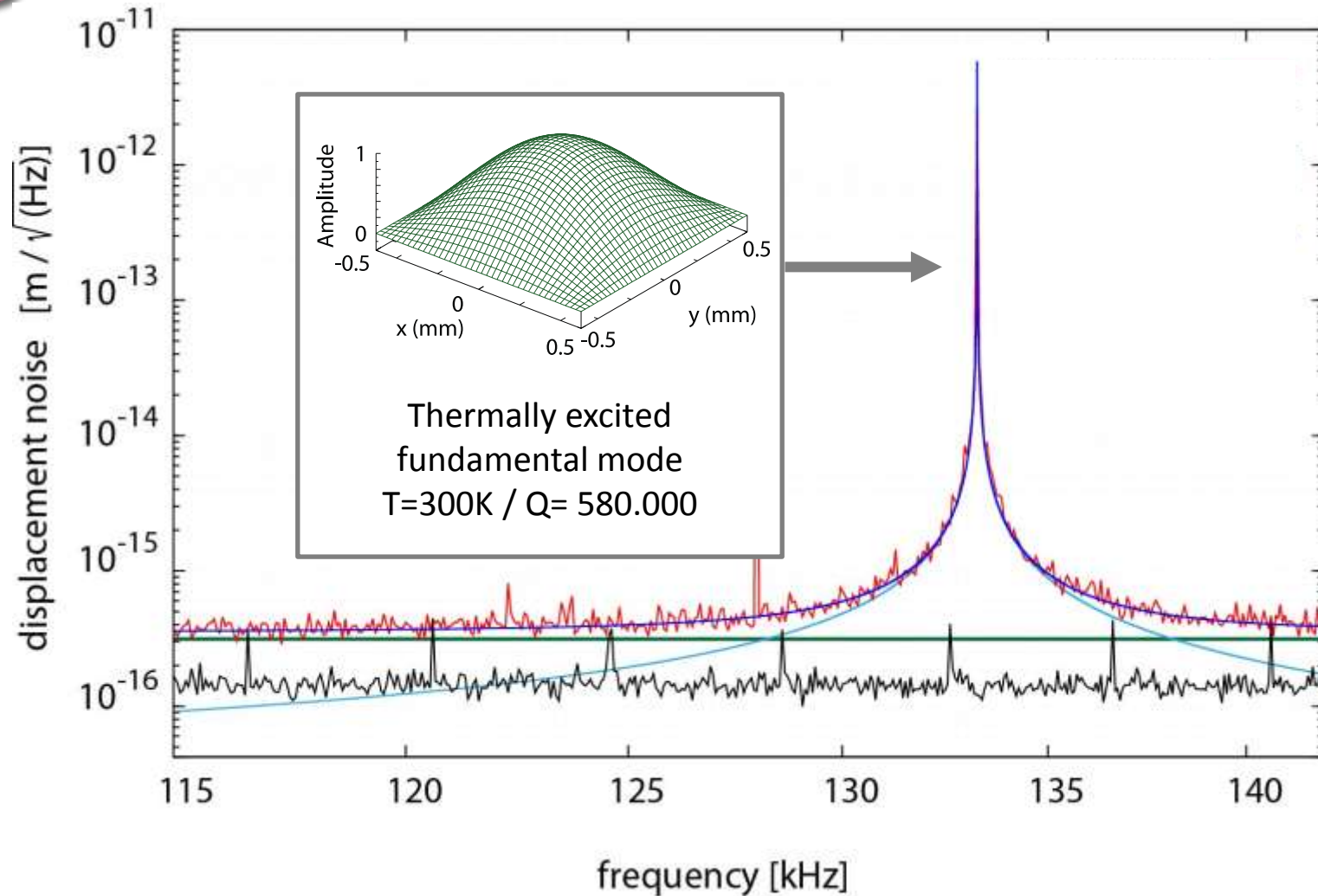
H. Kaufer et al. New J. Phys. **14** 095018 (2012)

MiSa-Interferometer as a mirror

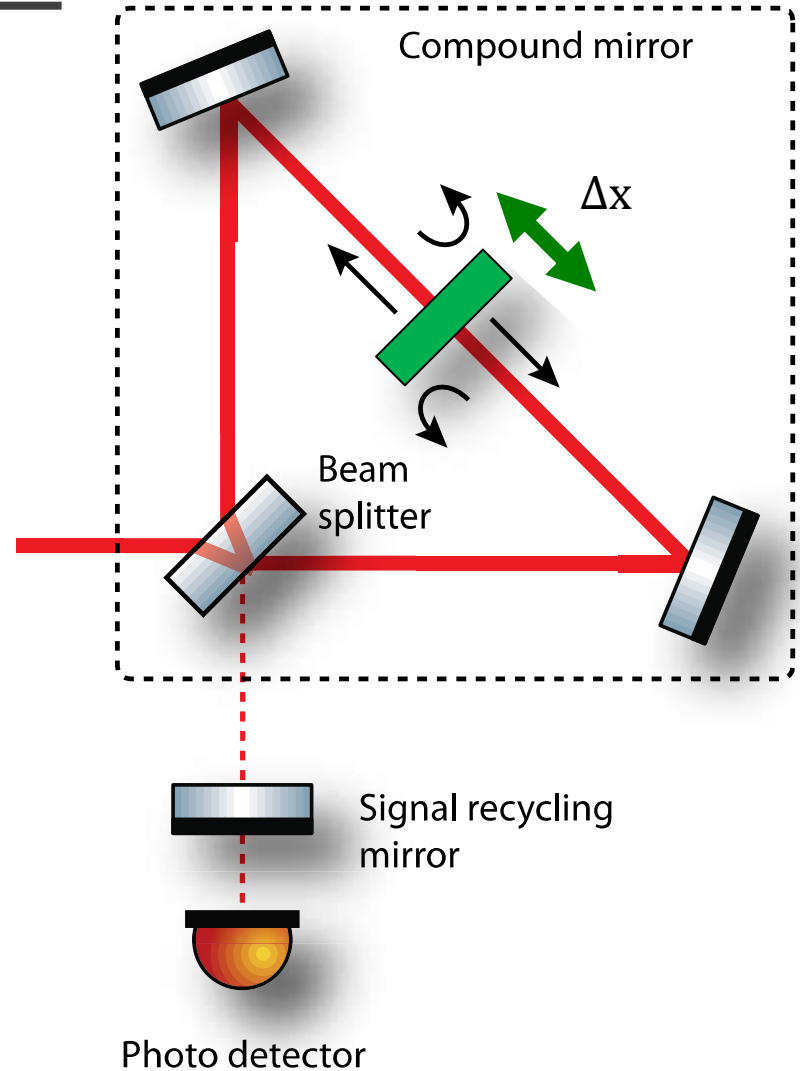
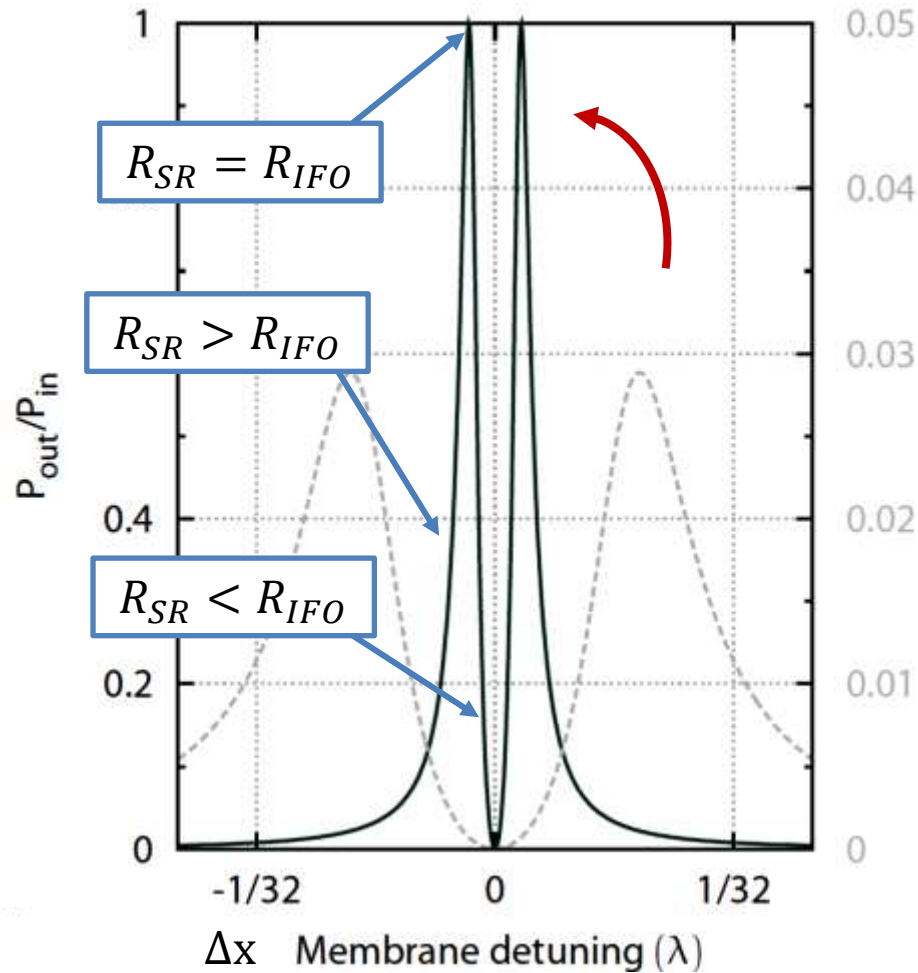




MiSa-Interferometer Output Spectrum

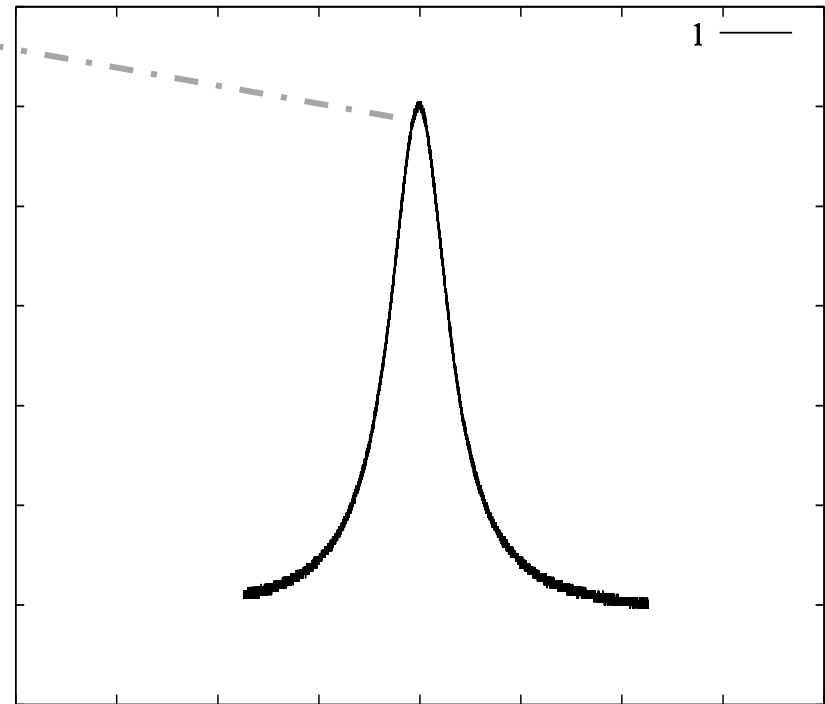
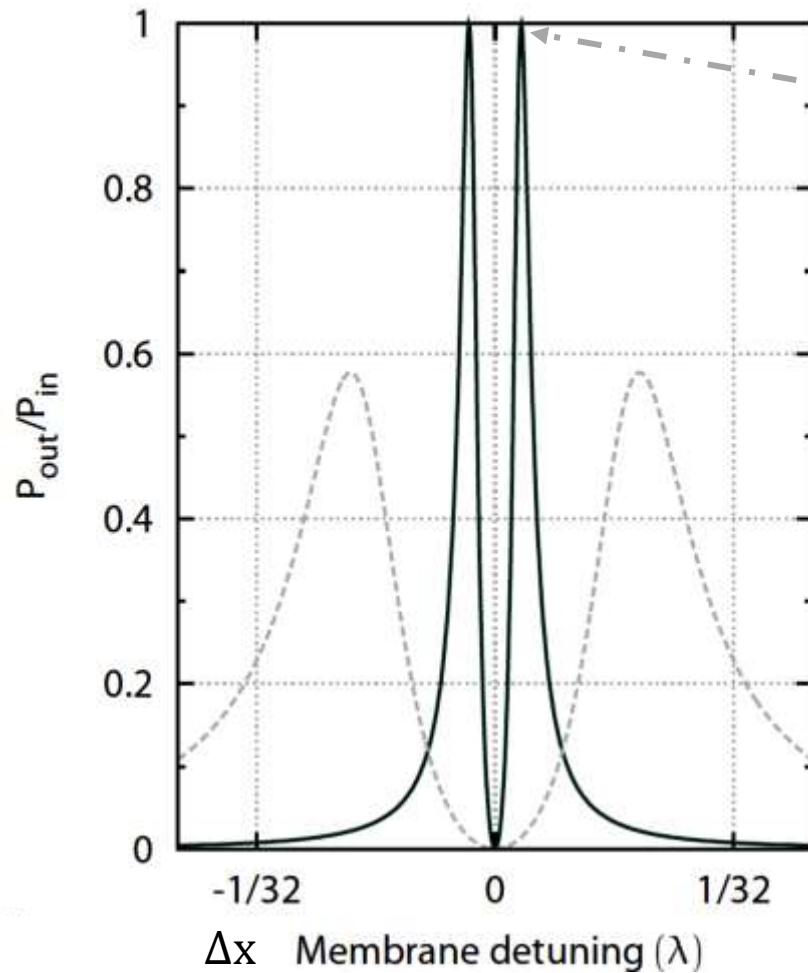


Signal Recycling Cavity



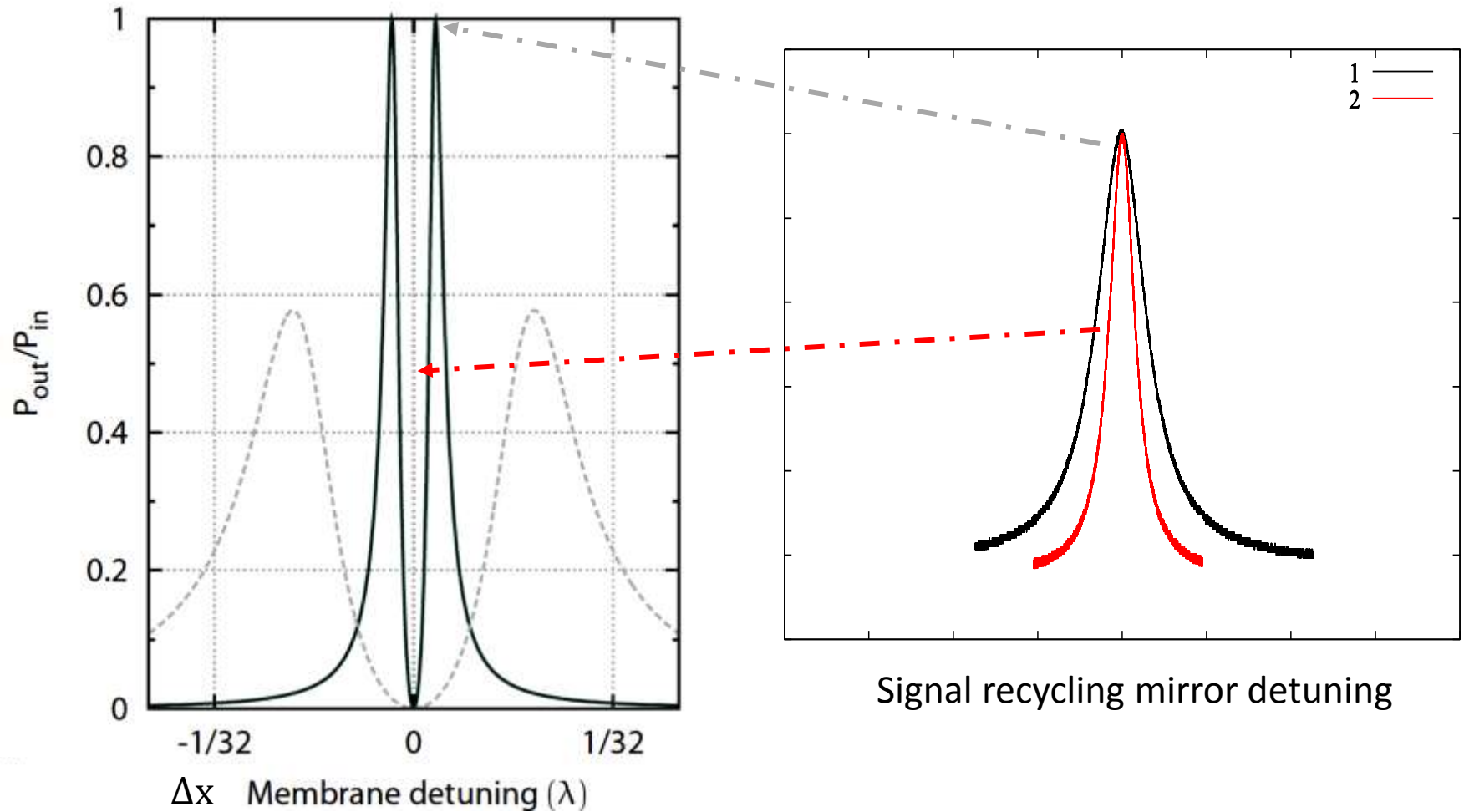


Linewidth depends on membrane position



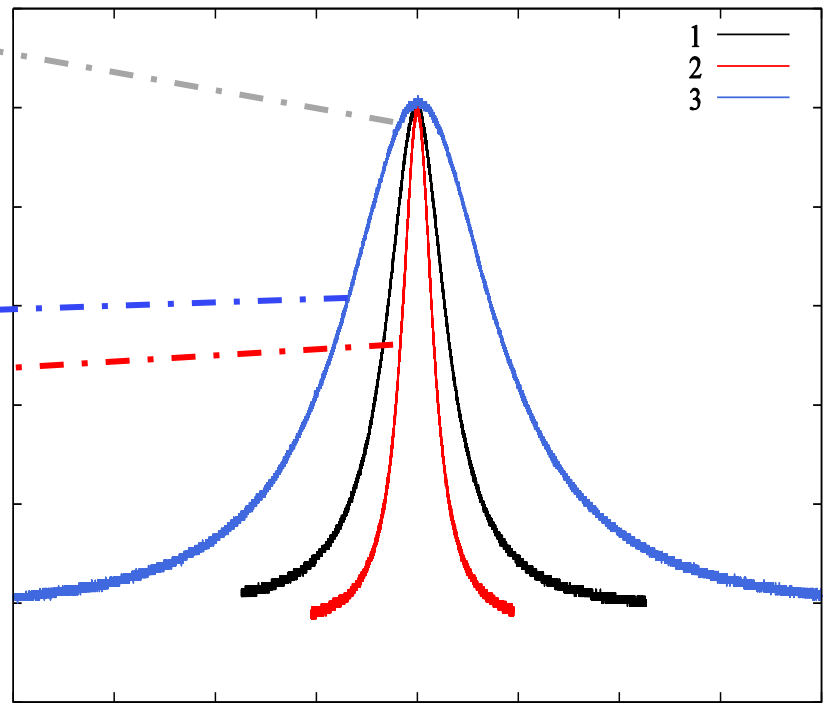
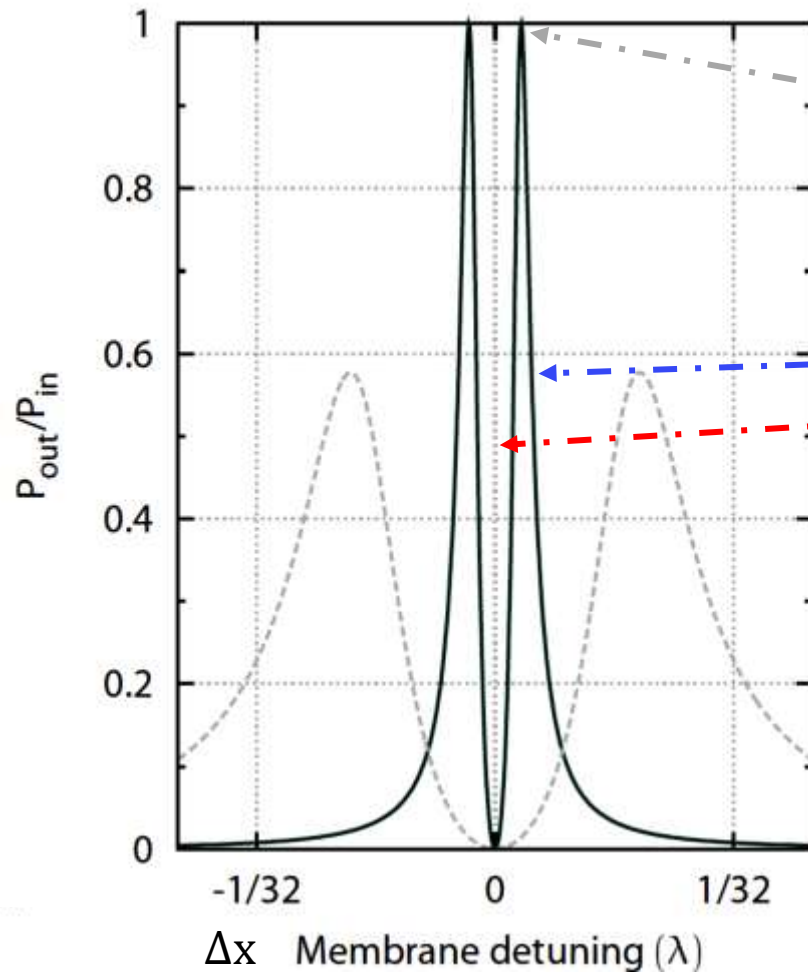
Signal recycling mirror detuning

Linewidth depends on membrane position





Linewidth depends on membrane position



Signal recycling mirror detuning

Dissipative Coupling

Modulation of the
cavity linewidth



Changing the coupling of the
optical system to the “thermal bath”

Interaction Hamiltonian:

$$\mathcal{H}_{int} = g_{\omega} \hat{x} \hat{a}^+ \hat{a} + g_{\gamma} \hat{x} \dots$$

dispersive coupling
(modulation of
cavity frequency)

$$g_{\omega} = \frac{\delta\omega}{\delta x}$$

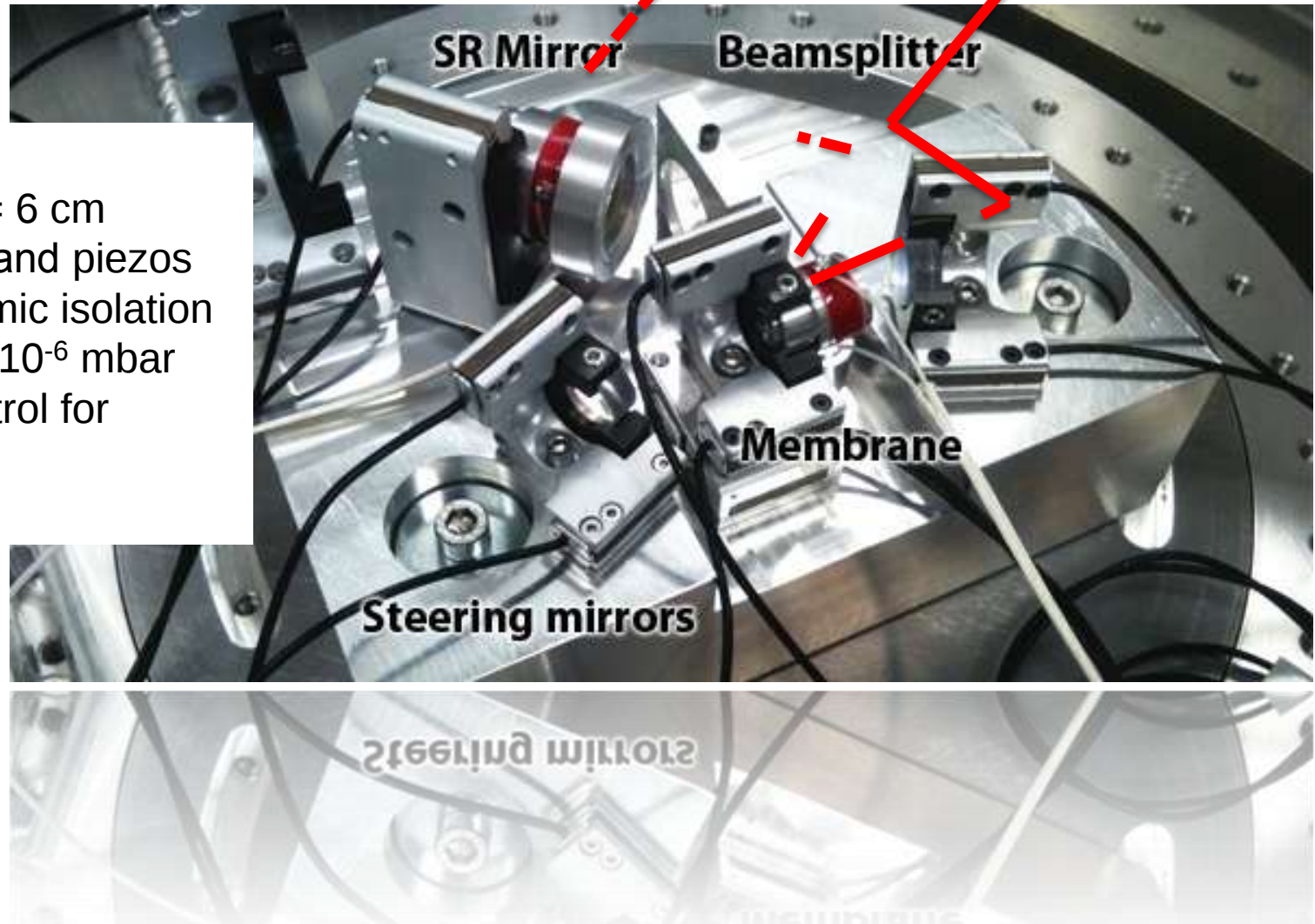
dissipative coupling
(modulation linewidth)

$$g_{\gamma} = \frac{\delta\gamma}{\delta x}$$

*F. Elste, S. Girvin, A. Clerk, PRL **102**, 207209, 2009*



Experimental Setup



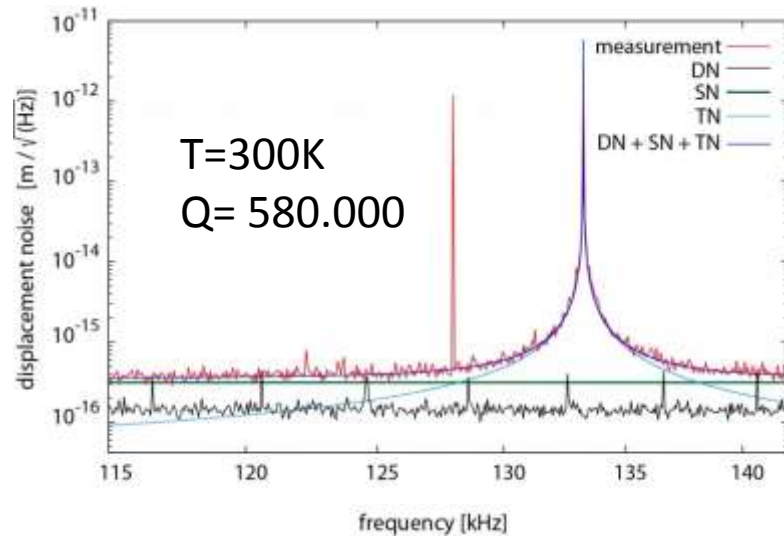
Setup features

- Arm length = 6 cm
- 0.5 " Optics and piezos
- Double seismic isolation
- Operated at 10^{-6} mbar
- Remote control for alignment

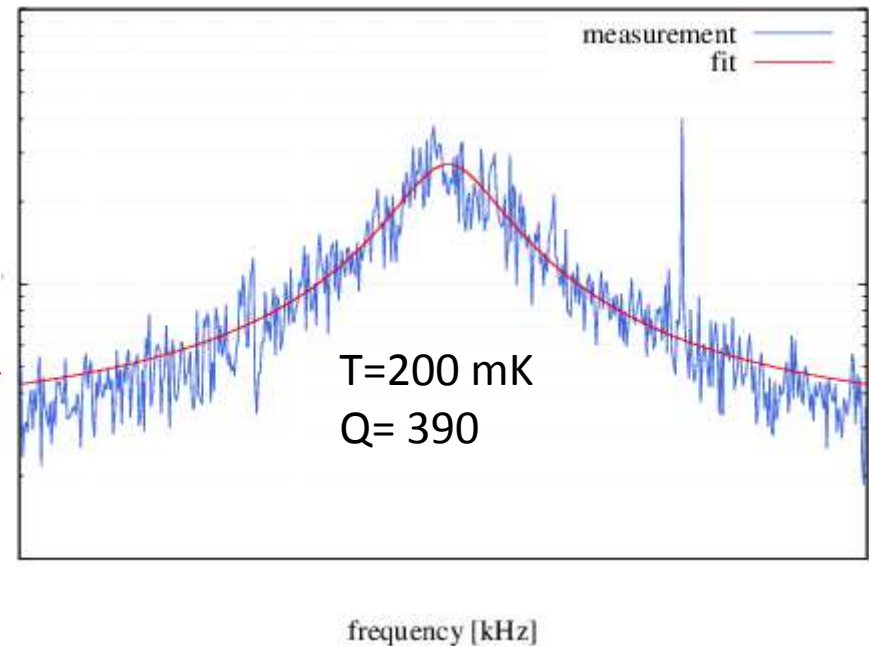


Cooling of the membrane

without signal- recycling

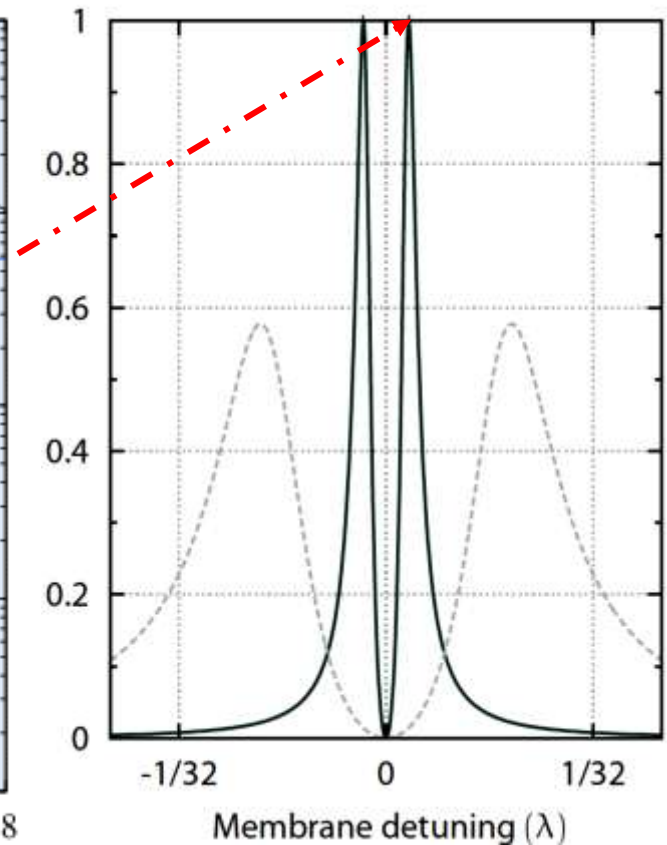
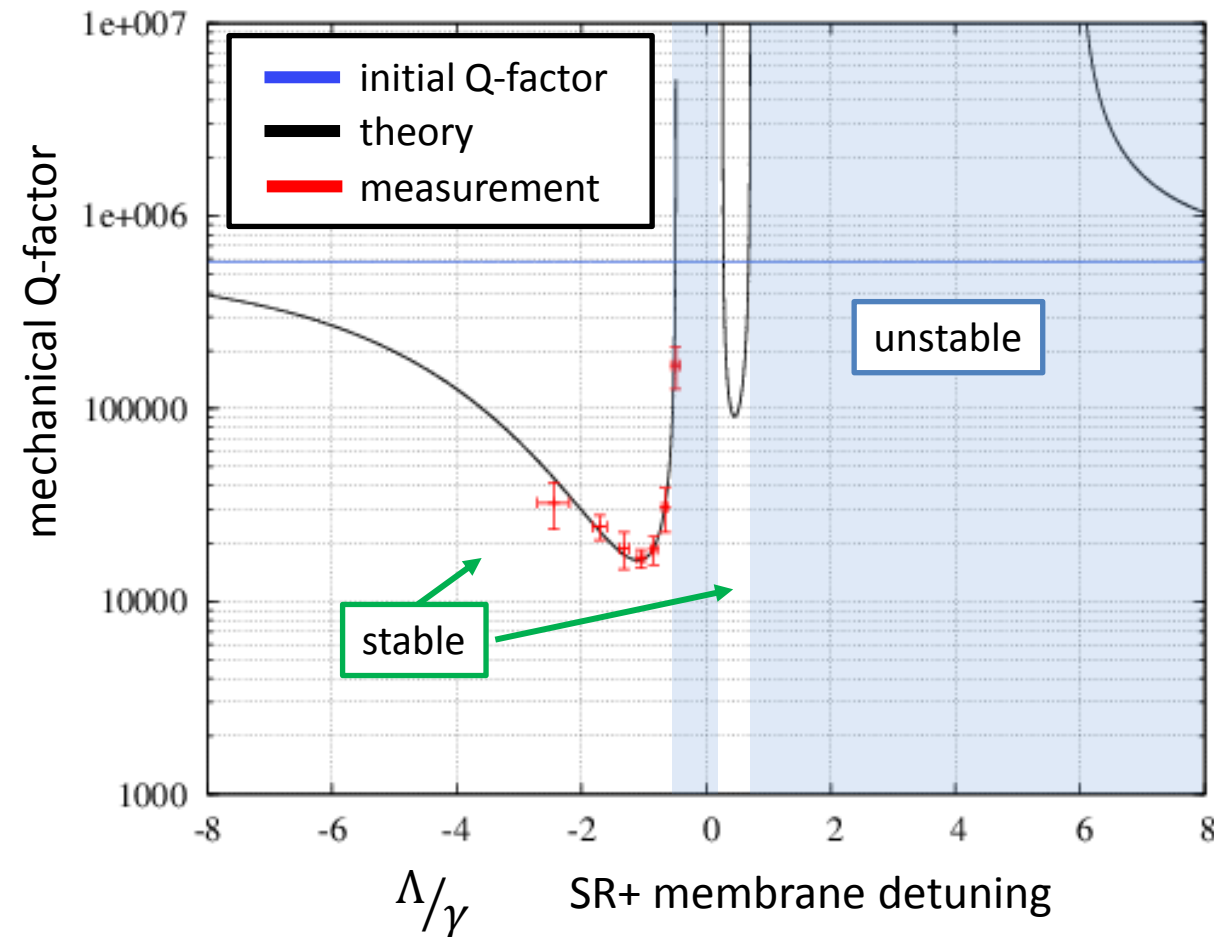


with signal- recycling



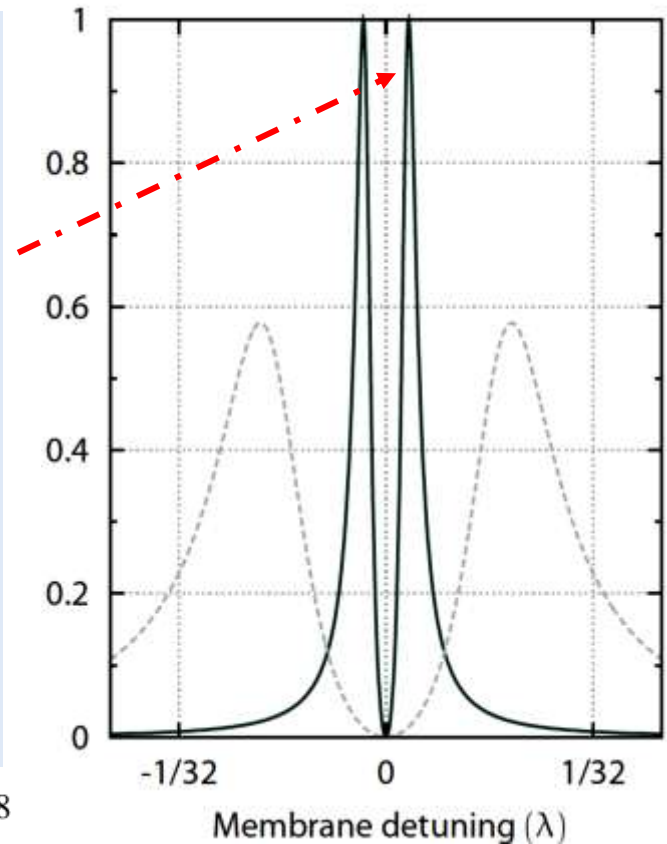
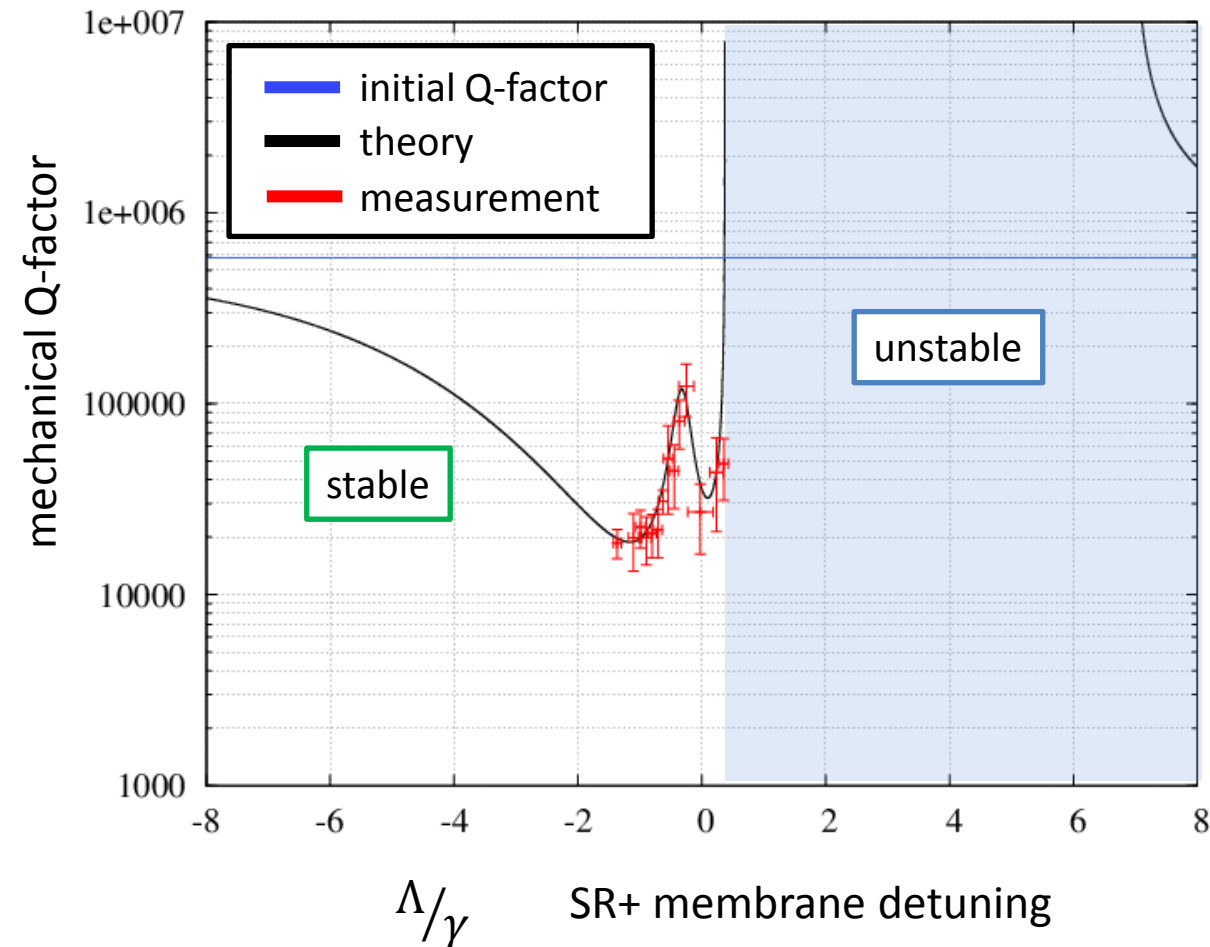


Stable + unstable regime measurements



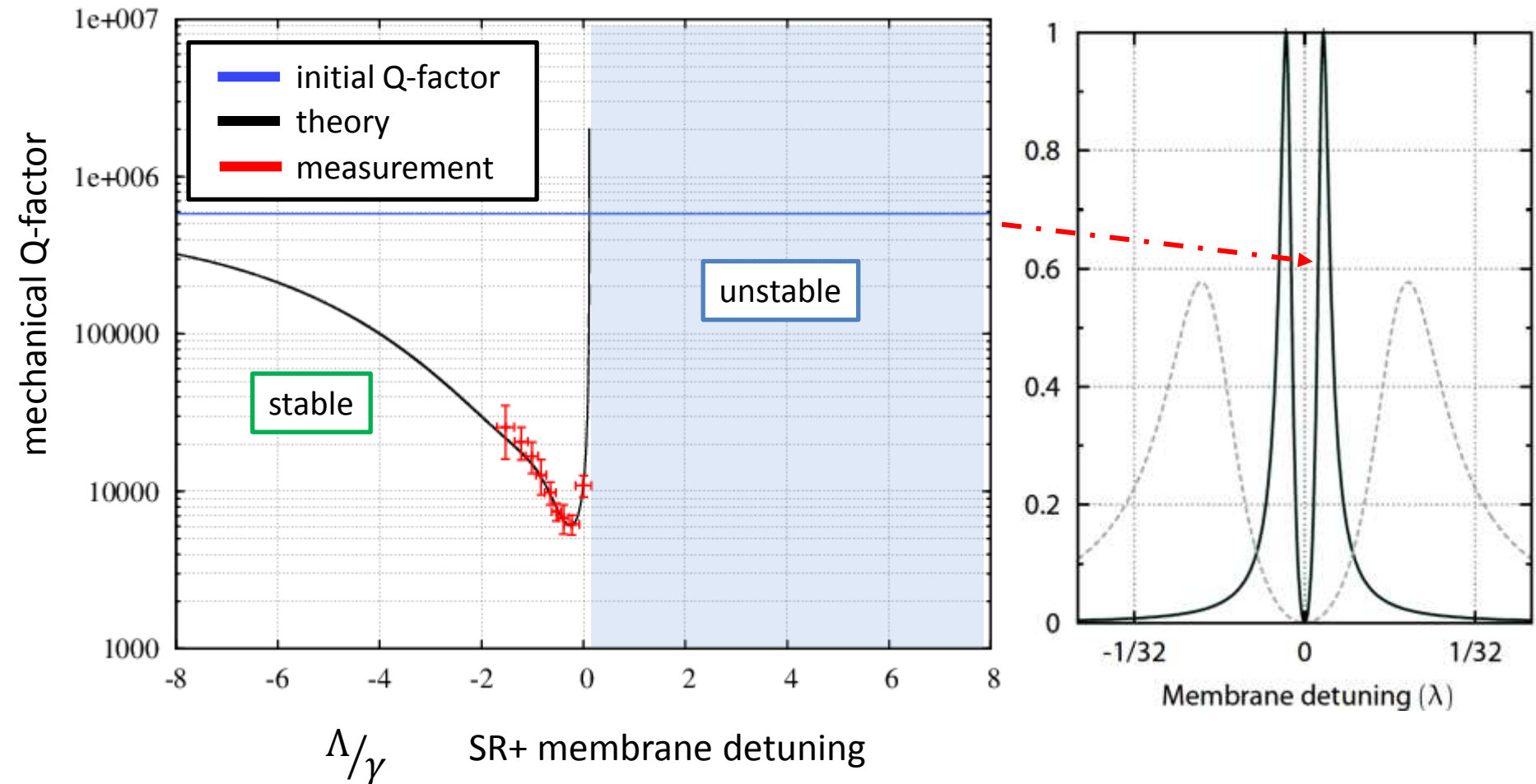


Stable + unstable regime measurements





Stable + unstable regime measurements



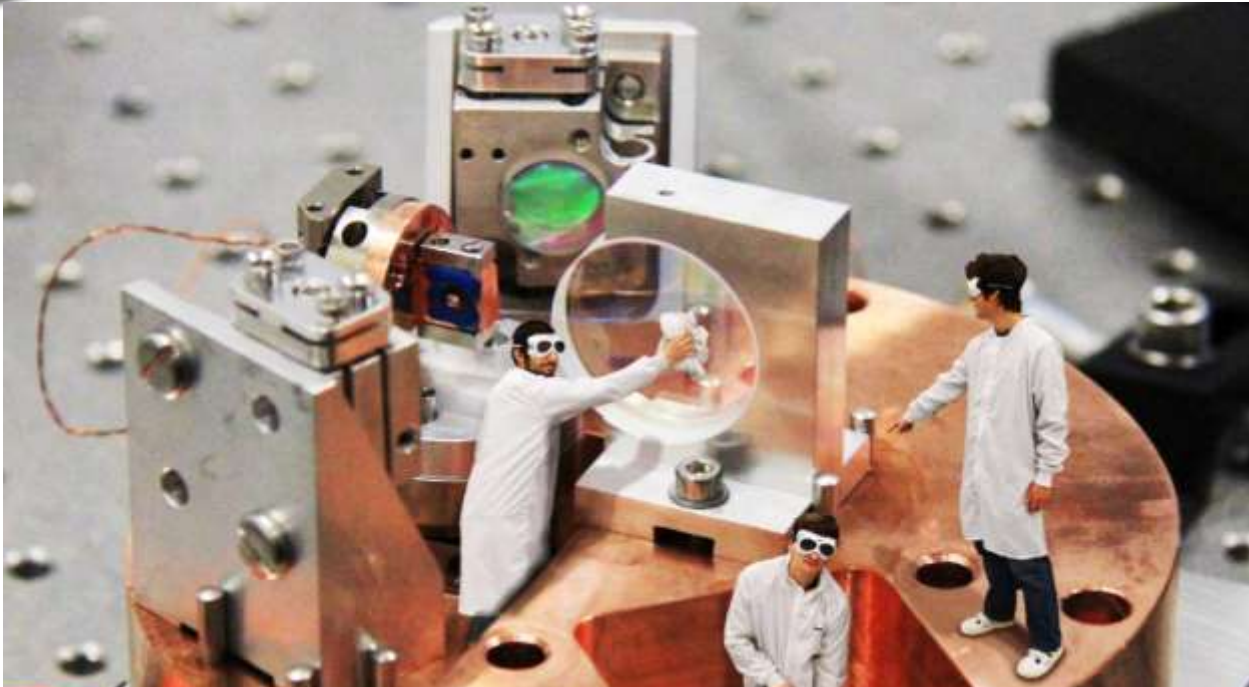


Conclusion

- ➔ Dynamical back-action off dark-port investigated and theory by Tarabrin confirmed
- ➔ Optical spring have some stable regions.
Stable optical spring in GWD with single carrier. Usefull?
- ➔ Operation off dark-port leads to new type of dynamical back-action: dissipative coupling



optomechanics@aei



Henning
Kaufer
(former PHD)



Andreas
Sawadsky
PHD

τῆαηκ γομ !

